

Scalar Quantum Field Theory on AdS spacetimes: Boundary conditions and Hadamard states

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Dipartimento di Fisica – Università di Pavia



Outline of the Talk

- Scalar Fields on AdS Spacetimes
- Boundary Conditions and Quantization: Mode Expansion
- Unruh-de Witt Detector and anti-Hawking effect
- The Hadamard expansion

Based on

- C. D. and H. R. C. Ferreira, Phys. Rev. D **94** (2016) no.12, 125016 & Rev. Math. Phys. **30** (2017) no.02, 1850004
- C. D., N. Drago and H. Ferreira, Lett. Math. Phys. **109** (2019) no.10, 2157–2186
- L. De Souza Campos and C. D., Phys. Lett. B **816** (2021), 136198 & Phys. Rev. D **103** (2021) no.2, 025021
- B. Costeri, C. D., B. A. Juárez-Aubry and R. D. Singh, [arXiv:2509.26035 [math-ph]].



The Poincaré patch of AdS_{d+1}

In the Poincaré chart, PAdS_{d+1} , the metric reads

$$ds^2 = \frac{\ell^2}{z^2}[-dt^2 + dz^2 + \delta^{ij}dx_i dx_j], \quad i, j = 1, \dots, d-1$$

where $z > 0$. Observe that

- 1 PAdS_{d+1} is **conformally related** to the upper half plane (\mathbb{H}^{d+1}, η) with conformal factor $\Omega = \frac{z}{\ell}$.
- 2 The hyperplane $z = 0$ in \mathbb{R}^{d+1} is the conformal boundary of PAdS_{d+1} .



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AQFT - I : Scalar field

Goal: Outline AQFT via a good example!

Basic assumptions.

- (M, g) is an arbitrary $4D$ globally hyperbolic spacetime. Hence, up to an isometry, $M \simeq \mathbb{R} \times \Sigma$

$$ds^2 = -\beta dt^2 + h_t, \quad \beta \in C^\infty(M; \mathbb{R}^+) \text{ and } h_t \in \text{Riem}(\Sigma), \quad \forall t \in \mathbb{R}$$

- $\phi : M \rightarrow \mathbb{R}$ is a *conformally coupled real scalar field*

$$P\phi = \left(-\square + \frac{R}{6} \right) \phi = 0,$$

The space of all smooth solutions is

$$\mathcal{S}_P(M) = \{ \phi \in C^\infty(M) \mid \exists f \in C_{tc}^\infty(M) \text{ and } \phi = G(f) \},$$

where $G = G^+ - G^-$ is the causal propagator.



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AQFT - II : Useful facts

It will be useful later to keep in mind that

- $f \in C^\infty(M)$ is **timelike compact (tc)** iff

$\text{supp}(f) \cap J^+(p)$ and $\text{supp}(f) \cap J^-(p)$ is compact $\forall p \in M$.

- $G^\pm : C_{tc}^\infty(M) \rightarrow C^\infty(M)$ are the advanced (+) and retarded (−) fundamental solutions for P such that

- 1 $P \circ G^\pm = G^\pm \circ P = \text{id}_{C_{tc}^\infty(M)}$
- 2 $\forall f \in C_{tc}^\infty(M)$, it holds $\text{supp}(G^\pm(f)) \subseteq J^\mp(\text{supp}(f))$

- All *dynamical configurations* of a real scalar field are

$$\mathcal{S}_P(M) \simeq \frac{C_{tc}^\infty(M)}{P[C_{tc}^\infty(M)]}.$$



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AQFT - III : Classical Observables

Paradigm: A *classical observable* is an assignment of a real number to each dynamical configuration.

- for every $\alpha \in C_0^\infty(M)$, define $F_\alpha : C^\infty(M) \rightarrow \mathbb{R}$

$$F_\alpha(\phi) = (\alpha, \phi) \doteq \int_M d\mu_g \phi(x) \alpha(x),$$

- Implement *dynamics* at a dual level: For all $\phi \in \mathcal{S}_P(M)$

$$0 = F_\alpha(P\phi) = (\alpha, P\phi) = (P^*\alpha, \phi) = (P\alpha, \phi).$$

- We have built the following

$$\mathcal{V}[\alpha] \in \frac{C_0^\infty(M)}{P[C_0^\infty(M)]} \mapsto \mathcal{F}_{[\alpha]} : \mathcal{S}_P(M) \rightarrow \mathbb{R}$$

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AQFT - IV : Algebra of Observables

- 1 Construct the unital *Borchers-Uhlmann* $*$ -algebra

$$\mathcal{T}(M) = \bigoplus_{n=0}^{\infty} \mathcal{E}^{obs}(M)^{\otimes n},$$

where $\mathcal{E}^{obs}(M)^0 \doteq \mathbb{C}$ and the $*$ -operation is complex conjugation.

- 2 Construct the ideal $\mathcal{I}(M) \subset \mathcal{T}(M)$ generated by elements of the form

$$[\alpha] \otimes [\alpha'] - [\alpha'] \otimes [\alpha] - iG([\alpha], [\alpha'])\mathbb{I} \quad (\text{CCR})$$

where \mathbb{I} is the unit in $\mathcal{T}(M)$ and

$$G([\alpha], [\alpha']) \doteq (\alpha, G(\alpha')) = \int_M d\mu_g(x) \alpha(x) G(\alpha')(x).$$

- 3 Define the **Algebra of Fields**

$$\mathcal{F}(M) \doteq \frac{\mathcal{T}(M)}{\mathcal{I}(M)}.$$



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AQFT - V : States

We need still the notion of an algebraic quantum state

For any unital \ast -algebra \mathcal{A} , a state is a linear map $\omega : \mathcal{A} \rightarrow \mathbb{C}$

$$\omega(\mathbb{I}) = 1 \text{ (normal.)} \quad \text{and} \quad \omega(a^* a) \geq 0, \quad \forall a \in \mathcal{A} \text{ (posit.)},$$

where e is the unit in \mathcal{A} .

We recover the probabilistic interpretation of quantum theories via:

- **GNS theorem:** $(\omega, \mathcal{A}) \mapsto (\mathcal{D}_\omega, \pi_\omega, \Omega_\omega)$
 - \mathcal{D}_ω is a dense subspace of a Hilbert space \mathcal{H}_ω .
 - $\pi_\omega : \mathcal{A} \rightarrow \mathcal{L}(\mathcal{D}_\omega)$ is a representation
 - $\Omega_\omega \in \mathcal{D}_\omega$ such that $\|\Omega_\omega\| = 1$ and $\mathcal{H}_\omega = \overline{\pi_\omega(\mathcal{A})\Omega_\omega}$.

$$\omega(a) = (\Omega_\omega, \pi_\omega(a)\Omega_\omega)_{\mathcal{H}_\omega}.$$



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AQFT - VI : Hadamard States - why?

Are all states physically acceptable?

Not in the slightest

- Minimal requirements are:
 - existence of a mathematically well-behaved, covariant notion of Wick polynomials to deal with interactions,
 - same UV behaviour of the Minkowski vacuum ,
 - quantum fluctuations of all observables finite.

Answer: Hadamard States

- Add-on: Invariance under the action of all isometries.



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AQFT - VII : Hadamard States - how?

How do we characterize Hadamard states?

- Notice that choosing a state $\omega : \mathcal{F}(M) \rightarrow \mathbb{C}$ is equivalent to assigning

$$\omega_n(\alpha_1, \dots, \alpha_n), \forall n \in \mathbb{N} \text{ and } \forall \alpha_i \in C_0^\infty(M)$$

so that ω is normalized, positive and accounts for the CCRs and dynamics.

- A special class of states are **quasi-free/Gaussian**, i.e., for all $n \in \mathbb{N}$

$$\omega_{2n+1} = 0, \quad \omega_{2n}(\alpha_1, \dots, \alpha_n) = \sum_{\pi_{2n} \in S'_{2n}} \prod_{i=1}^n \omega_2(\alpha_{\pi_{2n}(i-1)}, \alpha_{\pi_{2n}(i)}) .$$

ω is constructed out of the two-point function $\omega_2 \in \mathcal{D}'(M \times M)$.

Definition

A quasi-free state ω is called **Hadamard** (Radzikowski (1996)) iff

$$WF(\omega_2) = \left\{ (x, y, k_x, -k_y) \in T^*M^2 \setminus \{0\} \mid (x, k_x) \sim (y, k_y), k_x \triangleright 0 \right\} .$$



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Building Hadamard states

How many Hadamard states do we know?

- **Existence:** Known for free fields via deformation arguments (S. A. Fulling, F. J. Narcowich and R. M. Wald, (1981)) or on static spacetimes (Sahlmann & Verch - 2000)
- **Explicit construction:** Known in highly symmetric spacetimes, e.g., Bunch Davies state in de Sitter or states for scalar fields on FRW (Olbermann - (2007), Them & Brum (2013), Degner (2013), Brum & Fredenhagen (2014))
- **Functional Analytic method:** Exploit pseudodifferential calculus:
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Exploit the asymptotic structure.



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Exploit **the asymptotic structure.**



Klein-Gordon field in PAdS_{d+1}

Consider $\phi : \text{PAdS}_{d+1} \rightarrow \mathbb{R}$

$$P\phi = (\square_{\text{PAdS}} - m_0^2 - \xi R)\phi = 0 \quad \xi \in \mathbb{R} \text{ and } R = -d(d+1)$$

- *Conformal rescaling* $\longrightarrow \Phi \doteq \Omega^{\frac{1-d}{2}} \phi : \mathbb{H}^{d+1} \rightarrow \mathbb{R}$ obeys

$$P_\eta \Phi \doteq \left(\square_\eta - \frac{m^2}{z^2} \right) \Phi = 0, \quad m^2 = m_0^2 + \left(\xi - \frac{d-1}{4d} R \right).$$

- We can expand

$$\Phi(\underline{x}, z) = \int_{\mathbb{R}^d} \frac{d^d \underline{k}}{(2\pi)^{\frac{d}{2}}} e^{i\mathbf{k} \cdot \mathbf{x}} \widehat{\Phi}_{\underline{k}}(z)$$

yields $(\lambda \doteq \omega^2 - \sum_{i=1}^{d-1} k_i^2)$

$$P_\eta \Phi = 0 \iff L\widehat{\Phi}_{\underline{k}} = \left(-\frac{d^2}{dz^2} + \frac{m^2}{z^2} - \lambda \right) \widehat{\Phi}_{\underline{k}} = 0.$$

This is a **singular Sturm-Liouville equation** on $(0, \infty)$.



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$$P\phi = (\square_{\text{PAdS}} - m_0^2 - \xi R)\phi = 0 \quad \xi \in \mathbb{R} \text{ and } R = -d(d+1)$$

- *Conformal rescaling* $\longrightarrow \Phi \doteq \Omega^{\frac{1-d}{2}} \phi : \mathring{\mathbb{H}}^{d+1} \rightarrow \mathbb{R}$ obeys

$$P_\eta \Phi \doteq \left(\square_\eta - \frac{m^2}{z^2} \right) \Phi = 0, \quad m^2 = m_0^2 + \left(\xi - \frac{d-1}{4d} R \right).$$

- We can expand

$$\Phi(\underline{x}, z) = \int_{\mathbb{R}^d} \frac{d^d \underline{k}}{(2\pi)^{\frac{d}{2}}} e^{i \underline{k} \cdot \underline{x}} \widehat{\Phi}_{\underline{k}}(z)$$

yields $(\lambda \doteq \omega^2 - \sum_{i=1}^{d-1} k_i^2)$

$$P_\eta \Phi = 0 \iff L \widehat{\Phi}_{\underline{k}} = \left(-\frac{d^2}{dz^2} + \frac{m^2}{z^2} - \lambda \right) \widehat{\Phi}_{\underline{k}} = 0.$$

This is a **singular Sturm-Liouville equation** on $(0, \infty)$.



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The Endpoint Classification - I

The most general solution of $L\hat{\Phi}_{\underline{k}} = \lambda\hat{\Phi}_{\underline{k}}$ is for $\lambda > 0$

$$\hat{\Phi}_{\underline{k}}(z) = a(\underline{k})\sqrt{z}J_{\nu}(\sqrt{\lambda}z) + b(\underline{k})\sqrt{z}Y_{\nu}(\sqrt{\lambda}z),$$

where $\nu = \frac{1}{2}\sqrt{1+4m^2} \geq 0 \dots m^2 \in [-\frac{1}{4}, \infty)$, the BF bound.

Which boundary conditions are allowed? How do we implement them?

- $\sqrt{z}J_{\nu}(\sqrt{\lambda}z) \propto_{z \rightarrow 0} z^{\nu+\frac{1}{2}}$ and $\sqrt{z}Y_{\nu}(\sqrt{\lambda}z) \propto_{z \rightarrow 0} z^{-\nu+\frac{1}{2}}$

How to impose standard (Robin) boundary conditions at $z = 0$?

Singular Sturm-Liouville theory is the answer



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Boundary Conditions - I

Goal: Implement boundary conditions at $z = 0$

- 1 Choose $\Phi_1(z)$ the *principal solution* at $z = 0$, i.e. the unique one (up to scalar multiples)

$$\lim_{z \rightarrow 0} \frac{\Phi_1(z)}{\Phi(z)} = 0 \quad \forall \Phi(z) \mid L\Phi = \lambda\Phi, \lambda \in \mathbb{C}.$$

- 2 Pick a *second* L^2 -solution $\Phi_2(z)$, linearly indep. from Φ_1 (non unique)
- 3 Observe that, up to a scalar multiple, $\exists \alpha \in [0, \pi)$ such that

$$\Phi(z) = \cos \alpha \Phi_1(z) + \sin \alpha \Phi_2(z),$$

and that, for a regular endpoint at $z = 0$,

$$\cos \alpha \Phi(0) + \sin \alpha \Phi'(0) = 0 \iff \cos \alpha W_z[\Phi, \Phi_1] + \sin \alpha W_z[\Phi, \Phi_2] = 0$$

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Boundary Conditions - II

Recall that we consider $L = -\frac{d^2}{dz^2} + \frac{m^2}{z^2}$ and $\lambda = q^2 \doteq \omega^2 - \sum_{i=1}^{d-1} k_i^2$.

The fundamental pair of solutions $(\hat{\Phi}_{\underline{k}}^1, \hat{\Phi}_{\underline{k}}^2)$ is

$$\begin{aligned}\hat{\Phi}_{\underline{k}}^1(z) &= \sqrt{\frac{\pi}{2}} q^{-\nu} \sqrt{z} J_{\nu}(qz), \\ \hat{\Phi}_{\underline{k}}^2(z) &= \begin{cases} -\sqrt{\frac{\pi}{2}} q^{\nu} \sqrt{z} J_{-\nu}(qz), & \nu \in (0, 1), \\ -\sqrt{\frac{\pi}{2}} \sqrt{z} \left[Y_0(qz) - \frac{2}{\pi} \log(q) \right], & \nu = 0. \end{cases}\end{aligned}$$

$\nu = \frac{1}{2} \sqrt{1 + 4m^2}$	Classification of $z = 0$	Boundary condition at $z = 0$
$\nu = \frac{1}{2}$	Regular (R)	$\cot(\alpha) \hat{\Phi}_{\underline{k}}(0) + \hat{\Phi}_{\underline{k}}'(0) = 0$
$\nu \in [0, 1), \nu \neq \frac{1}{2}$	Limit-circle (LC)	$-\cot(\alpha) W_z[\hat{\Phi}_{\underline{k}}, \hat{\Phi}_{\underline{k}}^1] + W_z[\hat{\Phi}_{\underline{k}}, \hat{\Phi}_{\underline{k}}^2] = 0$
$\nu \in [1, \infty)$	Limit-point (LP)	Not required



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Ground States - Mode Expansion I

- 1 Let $\omega_{2,\mathbb{H}} \doteq (zz')^{\frac{1-d}{2}} \omega_2 \in \mathcal{D}'(\mathring{\mathbb{H}}^{d+1} \times \mathring{\mathbb{H}}^{d+1})$. It holds

$$(P_\eta \otimes \mathbb{I})\omega_{2,\mathbb{H}} = (\mathbb{I} \otimes P_\eta)\omega_{2,\mathbb{H}} = 0.$$

- 2 Consider the Fourier transform along $\mathbb{R}^d \ni \underline{x}$. Integral kernel:

$$\omega_{2,\mathbb{H}}(x, x') = \lim_{\epsilon \rightarrow 0^+} \int_0^\infty \frac{d\omega}{\sqrt{2\pi}} e^{i\omega(t-t'-i\epsilon)} \int_0^\infty dk \left(\frac{k}{r}\right)^{\frac{d-3}{2}} J_{\frac{d-3}{2}}(kr) \widehat{\omega}_{2,\underline{k}}(z, z').$$

with

$$(L \otimes \mathbb{I})\widehat{\omega}_{2,\underline{k}} = (\mathbb{I} \otimes L)\widehat{\omega}_{2,\underline{k}} = \lambda \widehat{\omega}_{2,\underline{k}}, \quad L = -\frac{d^2}{dz^2} + \frac{n^2}{z^2}.$$



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The case with $\nu \in (0, 1)$

In order to construct $\widehat{\omega}_{2,\underline{k}}$, we need that $iG(x, x') \doteq \omega_{2,\mathbb{H}}(x, x') - \omega_{2,\mathbb{H}}(x', x)$,

$$G(x, x')|_{t=t'} = 0, \quad \partial_t G(x, x') = -\partial_{t'} G(x, x')|_{t=t'} = \delta(x, x').$$

Setting $c = \cot \alpha$

$$\widehat{\omega}_{2,\underline{k}} = \sqrt{zz'} \int_0^\infty dq \, q \frac{[cJ_\nu(qz) - q^{2\nu} J_{-\nu}(qz)][cJ_\nu(qz') - q^{2\nu} J_{-\nu}(qz')]}{c^2 - 2cq^{2\nu} \cos(\pi\nu) + q^{4\nu}}, \quad c \leq 0$$

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- locally Hadamard and quasi-free/Gaussian



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Open Questions

Unknown Properties:

- 1 What is the most general class of allowed boundary conditions?
- 2 Setting $G^+ = \Theta(t - t')G$ and $G^- = -\Theta(t' - t)G$ is it true that

$$\text{supp}(G^\pm(f)) \subseteq J^\pm(\text{supp}(f)) \quad \forall f \in C_0^\infty(\mathring{\mathbb{H}}^{d+1}).$$

- 3 What are the singularities of $\omega_{2,\mathbb{H}}$:

$$\text{WF}(\omega_{2,\mathbb{H}}) = ?$$

- 4 Can we construct a local form for $\omega_{2,\mathbb{H}}$ and for G^\pm ?



Unruh-De Witt Detectors

Given $\Phi : M \rightarrow \mathbb{R}$ a real scalar field

Detector \longleftrightarrow 2 level system $\{|0\rangle, |\Omega\rangle\}$

$$H_{int} = c\chi(\tau)\Phi(x(\tau)) \otimes \mu(\tau), \quad c \in \mathbb{R} \text{ and } \chi \in C_0^\infty(\mathbb{R}),$$

- $x(\tau)$ is the detector **worldline** (stationary),
- $\mu(\tau) = |\Omega\rangle\langle 0|e^{i\Omega\tau} + |0\rangle\langle \Omega|e^{-i\Omega\tau}$ – **monopole-moment operator**

$$H = H_M \otimes \mathbb{I} + \mathbb{I} \otimes H_D + H_{int}$$



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$$H = H_\Phi \otimes \mathbb{I} + \mathbb{I} \otimes H_D + H_{int}.$$



Quantities of Interest

We are interested in the *detector response function*

$$\mathcal{F} = \int_{\tau_i}^{\tau_f} d\tau \int_{\tau_i}^{\tau_f} d\tau' e^{-i\Omega(\tau-\tau')} \chi(\tau) \chi(\tau') \omega_2(x(\tau) x'(\tau')),$$

where ω_2 is the two-point function of Φ .

Assuming *infinite interaction time*, i.e., $\chi(\tau) \rightarrow 1$ we look at

- detector response function

$$\int_{\mathbb{R}^2} d\tau d\tau' e^{-i\Omega(\tau-\tau')} \omega_2(x(\tau) x'(\tau')),$$

- transition rate

$$\dot{\mathcal{F}} = \int_{\mathbb{R}} ds e^{-i\Omega s} \omega_2(x(s)).$$



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Anti-Hawking effect¹

Consider a static black-hole spacetime $M = \mathbb{R} \times I \times \Sigma_j$

$$ds^2 = -f(r)dt^2 + h(r)dr^2 + r^2 d\Sigma_j(\theta_1, \dots, \theta_{n-2})$$

- 1 M has a *bifurcate horizon* with **surface gravity** κ_h ,
- 2 The detector measures a **local Hawking temperature** $T_D = \frac{\kappa_h}{2\pi\sqrt{f(r)}}$.
- 3 The response function of the detector satisfies *detailed balance* at T_D^{-1}

$$\frac{\dot{\mathcal{F}}(\Omega)}{\dot{\mathcal{F}}(-\Omega)} = e^{-\beta\sqrt{f(r)}\Omega}$$

$$\text{Anti-Hawking Effect} \iff \frac{\partial \mathcal{F}(T_D)}{\partial T_D} < 0$$

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Static BTZ - Geometry

It is a *3D black hole* spacetime:

$$ds^2 = - \left(\frac{r^2}{L^2} - M \right) dt^2 + \frac{dr^2}{\frac{r^2}{L^2} - M} + r^2 d\theta^2.$$

- It has a bifurcate Killing horizon at $r_H = L\sqrt{M}$,
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- It has a conformal *timelike boundary* at $r \rightarrow \infty$

Two-point functions can be constructed as in $PAdS_{d+1}$.



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Static BTZ - Two-point Function - I

Consider a massless, conformally coupled, real scalar field

$$\Phi : \text{BTZ} \rightarrow \mathbb{R}, \quad \left(\square_g - \frac{3}{4} \right) \Phi = 0.$$

We consider $z = \frac{r^2 - L^2 M}{L^2 M} \in (0, 1)$ and

$$\Phi(t, z, \theta) = \int_{\mathbb{R}} d\omega \sum_{l \in \mathbb{Z}} e^{i(\omega t + l\theta)} R_{\omega, l, \gamma}(z)$$

where

$$R_{\omega, l, \gamma}(z) = z^{\frac{i\omega}{2r_H}} (1-z)^{\frac{3}{4}} (\cos \gamma F_1(1-z) + \sin \gamma (1-z)^{-\frac{1}{2}} F_2(1-z))$$

where F_1, F_2 are hypergeometric functions depending on ω, M, l .



Static BTZ - Two-point Function - II

Ground and thermal states with **Robin boundary conditions** at $T = \frac{\sqrt{M}}{2\pi L}$

$$\omega_2(x, x') = \lim_{\epsilon \rightarrow 0^+} \sum_{l \in \mathbb{Z}} \int_{\mathbb{R}} \frac{d\omega}{2\pi} \Theta(\omega) e^{-i\omega(t-t'-i\epsilon)} \mathcal{N}_l R_{\omega, l, \gamma}(z) R_{\omega, l, \gamma}(z') e^{il(\theta-\theta')}$$

$$\omega_{2,T}(x, x') = \lim_{\epsilon \rightarrow 0^+} \sum_{l \in \mathbb{Z}} \int_{\mathbb{R}} \frac{d\omega}{2\pi} \Theta(\omega) \left[\frac{e^{\frac{\omega}{T}} e^{-i\omega(t-t'-i\epsilon)} + e^{i\omega(t-t'+i\epsilon)}}{e^{\frac{\omega}{T}} - 1} \right] \mathcal{N}_l R_{\omega, l, \gamma}(z) R_{\omega, l, \gamma}(z') e^{il(\theta-\theta')}$$

On a static trajectory $x(\tau) = (\tau, z_D, \theta_D)$ the **response function** reads for $\Omega < 0$

$$\dot{\mathcal{F}}_0 = \sum_{l \in \mathbb{Z}} \frac{\mathcal{N}_l}{2\pi} R_{\omega, l, \gamma}^2(z_D) \Big|_{\omega = -\sqrt{|g_{00}(z_D)|} |\Omega|}$$

and

$$\dot{\mathcal{F}} = \frac{\text{sign}(\Omega)}{e^{\text{sign}(\Omega) \frac{\omega}{T}} - 1} \Big|_{\omega = \sqrt{|g_{00}(z_D)|} |\Omega|} \dot{\mathcal{F}}_0.$$



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Static BTZ - Anti-Hawking Effect - l^2

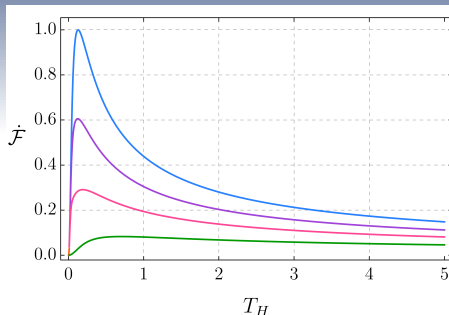


Figure: $l = 0$ contribution to $\dot{\mathcal{F}}_0$ as a function of T_H for $r_h = 1$, $\Omega = 0.1$ and different boundary conditions; from top to bottom, respectively, $\gamma = (0.50, 0.47, 0.40, 0.25, 0)\pi$.

²L. J. Henderson, R. A. Hennigar, R. B. Mann, A. R. H. Smith and J. Zhang, Phys. Lett. B **809** (2020), 135732



Static BTZ - Anti-Hawking Effect - II

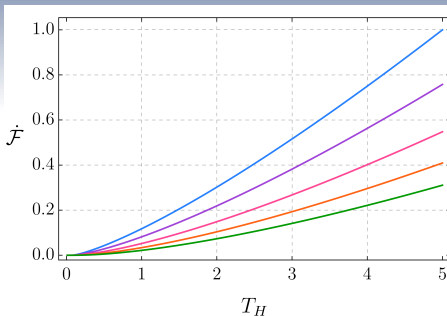


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Massless Hyperbolic Black Holes

These are n -dimensional spacetimes $M = \mathbb{R} \times I \times \Sigma_{n-2}$

$$ds^2 = -f(r)dt^2 + \frac{dr^2}{f(r)} + r^2 d\Sigma_{n-2}^2, \quad f(r) = -1 - \frac{2M}{r^{n-3}} + \frac{r^2}{L^2},$$

where $d\Sigma_{n-2}^2 = d\theta^2 + \sinh^2 \theta dS_{n-3}^2(\varphi_1, \dots, \varphi_{n-3})$.

- It has a bifurcate Killing horizon at r_H such that $f(r_H) = 0$,
- it is a solution of vacuum Einstein's equations with $\Lambda = -\frac{(n-1)(n-2)}{2L^2}$.
- It has a conformal *timelike boundary* at $r \rightarrow \infty$

Two-point functions can be constructed as in $PAdS_{d+1}$ and in BTZ.



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Anti-Hawking Effect - I

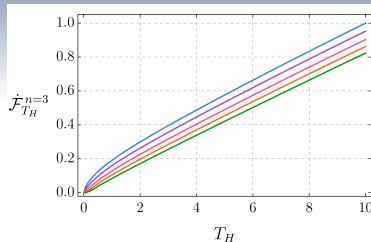
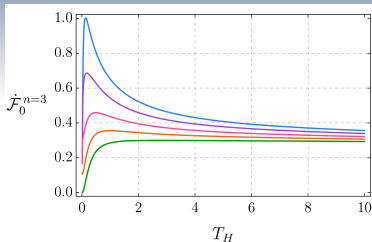


Figure: Transition rate, integrated up to $\ell = 100$, as a function of the local Hawking temperature on the three-dimensional hyperbolic black hole for $\Omega = -0.1$, $\theta = \pi^{-1}$ and for different boundary conditions. From top to bottom $\gamma = (0.50, 0.47, 0.40, 0.25, 0)\pi$. On the left, for the ground state; on the right, for the KMS state.



Anti-Hawking Effect - II

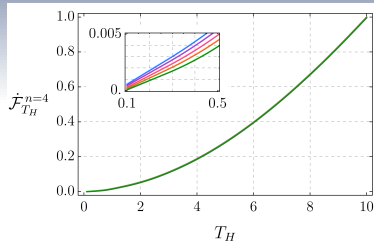
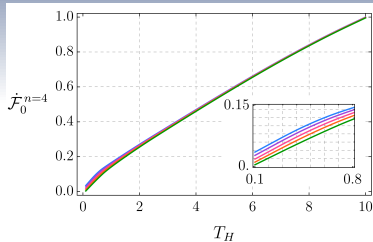


Figure: Transition rate as a function of the local Hawking temperature, summed up to $m_1 = 100$ and integrated up to $\ell = 100$, on the four-dimensional hyperbolic black hole for $\Omega = -0.1$, at $\theta_D = \pi^{-1}$, $\varphi_{1,D} = 0$ and for different boundary conditions: from top to bottom $\gamma = (0.50, 0.47, 0.40, 0.25, 0)\pi$. On the left, for the ground state; on the right, for the KMS state.



The Bondurant-Fulling map

Simplified setting

- $(M, g) \longrightarrow (\mathbb{H}^d, \eta)$ with $\mathbb{H}^d = \{(t, x_1, \dots, x_{d-2}, z) \mid z \geq 0\}$
- Massive scalar field $\Phi : \mathbb{H}^d \rightarrow \mathbb{R}$

$$P\Phi = (\square_\eta - m^2)\Phi = 0,$$

- **Robin boundary conditions**

$$\partial_z \Phi|_{z=0} = \alpha \Phi|_{z=0}, \quad \alpha \in \mathbb{R}.$$

- Relevant Spaces:

$$C_D^\infty(\mathbb{R}^d) = \{f \in C^\infty(\mathbb{R}^d) \mid f|_{z=0} = 0\},$$

$$C_\alpha^\infty(\mathbb{R}^d) = \{f \in C^\infty(\mathbb{R}) \mid \partial_z f|_{z=0} = \alpha f|_{z=0}\}.$$

$$T_\alpha : C_\alpha^\infty(\mathbb{R}^d) \rightarrow C_D^\infty(\mathbb{R}^d) \quad f \mapsto T_\alpha(f) \doteq \partial_z f - \alpha f.$$



Constructing the propagators

Method of images

- Calling $G^\pm(\underline{x} - \underline{x}', z - z')$ the propagators on (\mathbb{R}^d, η)

$$G_{D/N}^\pm(\underline{x} - \underline{x}', z, z') = G^\pm(\underline{x} - \underline{x}', z - z') \mp G^\pm(\underline{x} - \underline{x}', z + z')$$

Notice

$$\begin{aligned} \text{WF}(G_{D/N}^\pm) = \\ \{(x, k_x, x', k_{x'}) \in T^*(\mathcal{M} \times \mathcal{M}) \setminus \{0\} \mid (x, k_x) \sim_\pm (x', -k_{x'}), k_x \neq 0\} \cup \text{WF}(\delta_2), \end{aligned}$$



Constructing the propagators - II

Denote by $\mathcal{L}_\alpha(z) = \Theta(z)e^{-\alpha z}$

$$G_\alpha^\pm := (\mathcal{L}_\alpha \otimes \delta) \star G_D^\pm \circ (\mathbb{I} \otimes T_\alpha) = G_N + 2\kappa \mathcal{L}_\alpha \star G(\underline{x} - \underline{x}', z + z'),$$

- G_α^\pm coincide with the operator counterparts,
- $\text{supp}(G_\alpha^\pm(f)) \subseteq J^\mp(\text{supp}(f))$ for all $f \in C_0^\infty(\mathbb{H}^d)$.
- the wavefront set is

$$\begin{aligned} \text{WF}(G_\alpha^\pm) = \\ \{(x, k_x, x', k_{x'}) \in T^*(\mathcal{M} \times \mathcal{M}) \setminus \{0\} \mid (x, k_x) \sim_\pm (x', -k_{x'}), k_x \neq 0\} \cup \text{WF}(\delta_2), \end{aligned}$$



Propagators - Local Form

$$G_{\alpha}^{+}(x, x') = \Theta(t - t') \left(U(x, x') \delta^{\frac{d-2}{2}}(\sigma) + \delta_d V(x, x') \Theta(\sigma) + \right. \\ \left. + U'(x, x') \delta^{\frac{d-2}{2}}(\sigma_-) + \delta_d V'(x, x') \Theta(\sigma_-) \right)$$

where

$$\delta_d = \begin{cases} 1 & \text{if } d \text{ is even} \\ 0 & \text{if } d \text{ is odd} \end{cases} .$$



Propagators - Recursion Relations

In the even-dimensional case:

$$\begin{cases} U(x, x') = \sum_{j=0}^{\frac{d-4}{2}} u_j(\underline{x}, z, \underline{x}', z') \sigma^j \\ V(x, x') = \sum_{j=0}^{\infty} v_j(\underline{x}, z, \underline{x}', z') \sigma^j \\ U'(x, x') = \sum_{j=0}^{\frac{d-4}{2}} u'_j(\underline{x}, z, \underline{x}', z') \sigma^j_- \\ V'(x, x') = \sum_{j=0}^{\infty} v'_j(\underline{x}, z, \underline{x}', z') \sigma^j_- \end{cases},$$

where, imposing the defining equations for G_α^+

$$\begin{cases} (2-d)\sigma^\mu \partial_\mu u_0 = 0, \\ [u_0] = 1, \\ Pu_j + (2j+4-d)\sigma^\mu \partial_\mu u_{j+1} + (j+1)(2j+4-d)u_{j+1} = 0, \\ [u_{j+1}] = -\frac{[Pu_j]}{2j+4-d}, \quad 0 \leq j \leq \frac{d}{2} - 3, \end{cases}$$



Propagators - Recursion Relations 2

The reflected components abide by

$$\begin{cases} (2-d)\sigma_-^\mu \partial_\mu u'_0 = 0, \\ u'_0|_{z=0} = u_0|_{z=0}, \\ Pu'_j + (2j+4-d)\sigma_-^\mu \partial_\mu u'_{j+1} + (j+1)(2j+4-d)u'_{j+1} = 0, \\ (\partial_z + \kappa)(u_j + u'_j)|_{z=0} + \frac{1}{2}(2j+4-d)\partial_z \sigma(u_{j+1} - u'_{j+1})|_{z=0} = 0, \quad 0 \leq j \leq \frac{d}{2} - 3. \end{cases}$$

- Similar transport equations for v_j and v'_j ,
- The series are converging uniformly on compact subsets.



Global-to-Local Theorem

Theorem (Global-to-Local)

Given (\mathbb{H}^d, η) and $P := \square_\eta + m^2$, consider $\omega_{2,\kappa} \in \mathcal{D}'(\mathbb{H}^d \times \mathbb{H}^d)$ abiding by Robin boundary conditions with $\kappa \geq 0$. The following statements are equivalent:

1. $\omega_{2,\kappa}$ reads

$$\begin{aligned} \omega_{2,\kappa}(x, x') &= \lim_{\epsilon \rightarrow 0^+} U(x, x') \sigma_\epsilon^{\frac{2-d}{2}} + \delta_d V(x, x') \ln \left(\frac{\sigma_\epsilon}{\lambda^2} \right) + \\ &U'(x, x') \sigma_{-, \epsilon}^{\frac{2-d}{2}} + \delta_d V'(x, x') \ln \left(\frac{\sigma_{-, \epsilon}}{\lambda^2} \right) + W(x, x'), \end{aligned}$$

2. $\omega_{2,\kappa}$ has the following singular structure:

$$\begin{aligned} WF(\omega_{2,\kappa}) &= \{(x, k, x', -k') \in T^*(\mathring{\mathbb{H}}^d \times \mathring{\mathbb{H}}^d) \setminus \{\mathbf{0}\} \mid \\ &(x, k) \sim (x', k') \text{ and } k \triangleright 0\}, \end{aligned}$$



Conclusions

Conclusions:

- Construction of fundamental solutions and states in PAdS_{d+1} ,
- Identification of admissible boundary conditions,
- Application to black-hole physics,
- Characterization of the local form.

Open Questions:

- Generalization of the local form to PAdS_{d+1} ,
- Application of the framework to interacting field theories.