

XENONnT Experiment

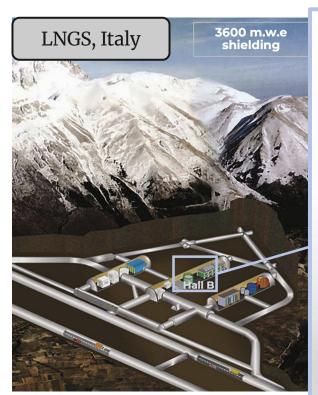
Direct detection of dark matter

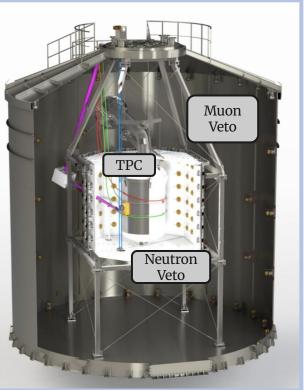
200 +**Scientists**

30 **Institutions**

12 Countries







Main Objective:

Discover Weakly Interacting Massive Particles (WIMPs)

XENON program timeline







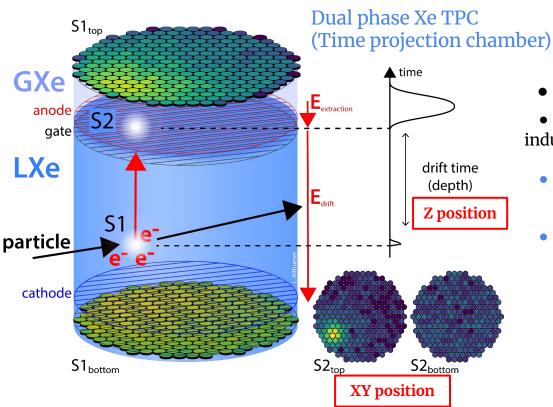




XENONnT Experiment

Direct detection of dark matter





- **S1**: Prompt scintillation light
- **S2:** Secondary scintilation light induced by ionionized electrons
 - 3D Position reconstruction: Drift time + PMT pattern
 - Energy reconstruction:

$$E = W(\frac{cS_1}{g_1} + \frac{cS_2}{g_2})$$

W: Average energy to produce a quanta cS1, cS2: Corrected area of S1 and S2

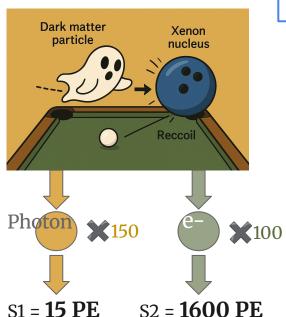
g1, g2: Gain of S1 and S2

Lower the detection threshold \Rightarrow Light DM



Why S2-only? Small example:

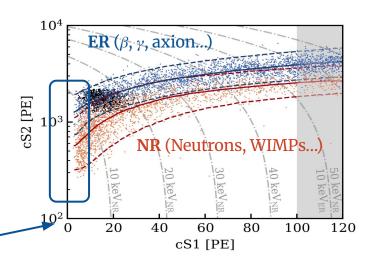
Q~2.8 keV from Ar37 decay



Detected Photoelectron [PE]

S2 amplified by secondary scintillation

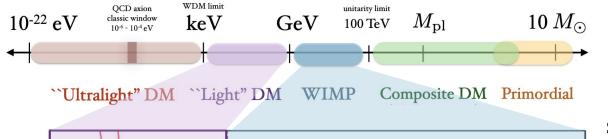


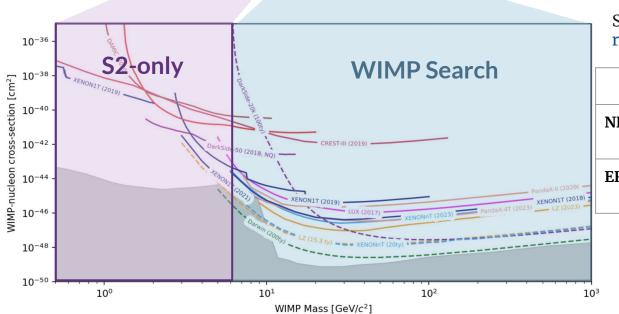


NR: Nuclear recoil ER: Eletronic recoil

Lower the detection threshold \Rightarrow Light DM







S2-only lower the detected recoiled energy:

| | S1-S2 | S2-only |
|----|-----------------------|-----------------------------|
| NR | 3.5 keV (6 GeV DM) | 0.7 keV (3 GeV DM) |
| ER | 1.2 keV | 0.18 keV (0.05-3 GeV DM) |

 10^{-16}

(e) Dark Photon

DM Mass [keV]

 10^{-2}

Dark matter model and recent results

Is done:

(f) Axion-like Particles

0.1

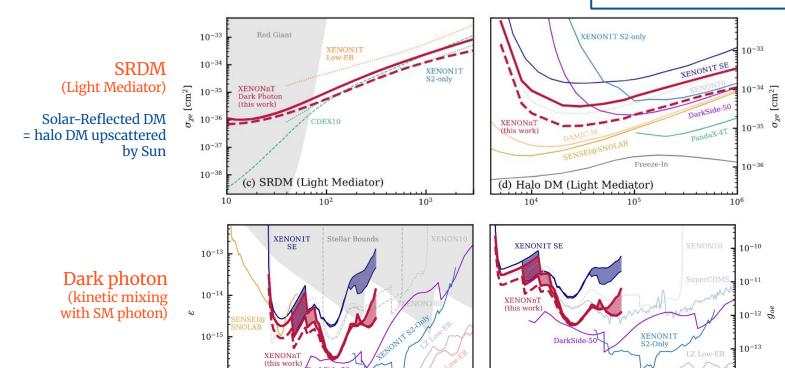
DM Mass [keV]

10

- without background model
- < 5 detected electrons

Better results with background model?

 10^{-14}



Halo DM (Light Mediator)

With velocity distribution from standard halo model

Axion-like particles (electron coupling)

S2-only analysis (current work)

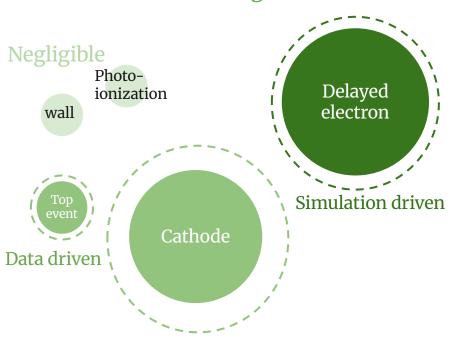
Major challenge: Background model

Without S1 \Rightarrow No z-related data selection

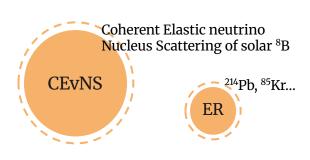
Small S2 area \Rightarrow More sensitive to detector condition



Instrumental background



Physics background







S2-only analysis (current work)

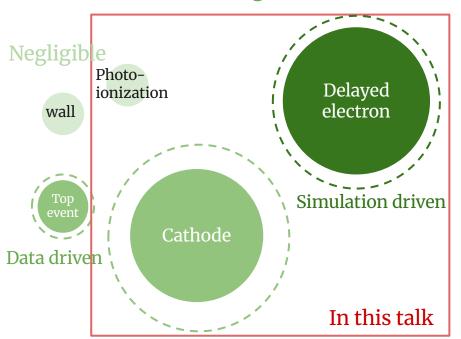
Major challenge: Background model

Without S1 \Rightarrow No z-related data selection

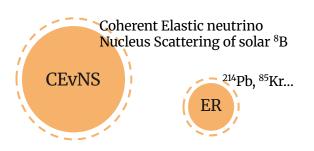
Small S2 area \Rightarrow More sensitive to detector condition



Instrumental background



Physics background



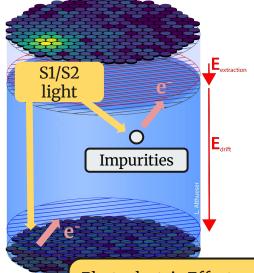




Major background: Photoionization vs Delayed electrons



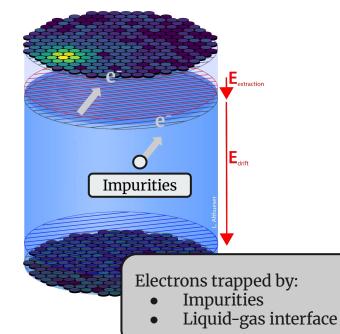
Photoionization



Photoelectric Effect:

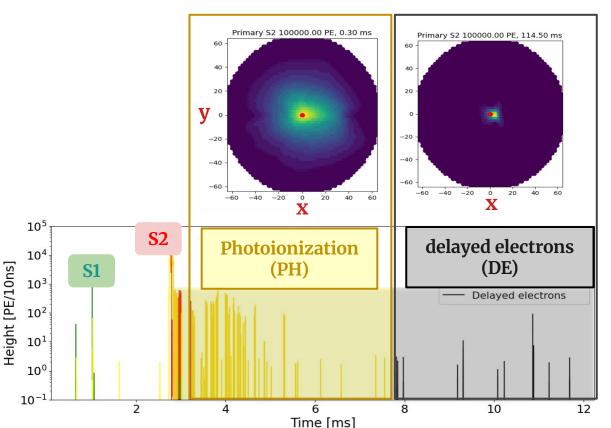
- Impurities in the liquid xenon Metal surfaces

Delayed electrons



Major background: Photoionization vs Delayed electrons





Position distribution

PH: no strongly position correlated with previous S2

DE: strongly position correlated with previous S2

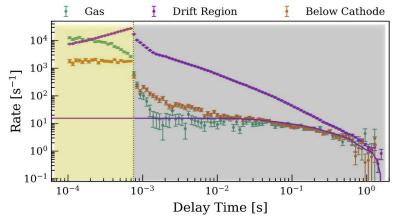
Time distribution

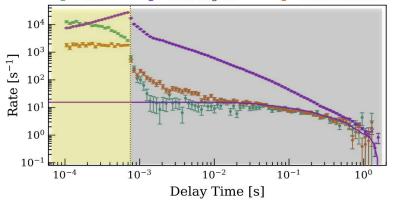
PH: within 2 * maximum drift time

DE: outside 2 * maximum drift time

Major background: Photoionization vs Delayed electrons







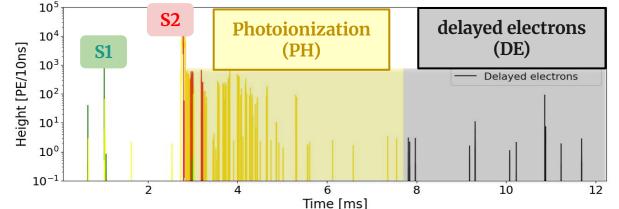


Can be identified and rejected in time:

Time veto cut

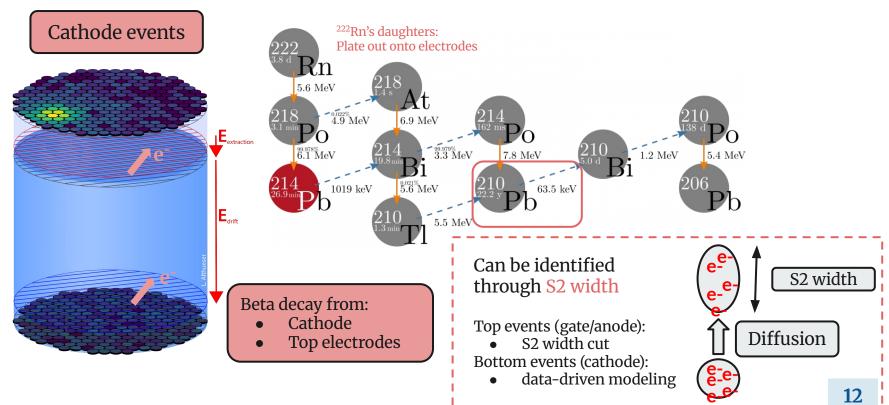
Delayed electron:

- Simulation-driven modeling
- Choose large S2
- Simulate following small S2
 - Area
 - Delayed time
 - Position

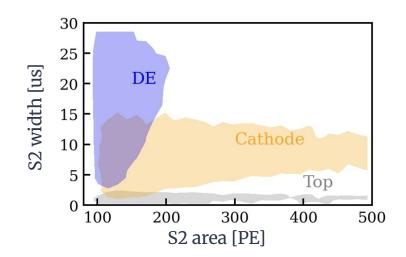


Major background: Cathode events



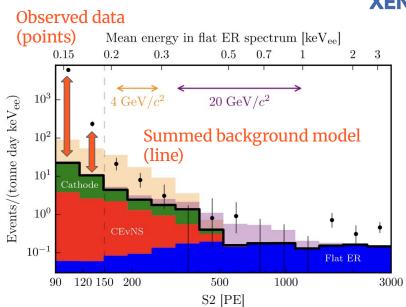


Future work: Background Validation



Current instrumental background estimation in XENONnT





Background validation in XENON1T

Summary

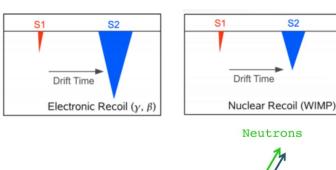
- XENONnT is a dual-phase xenon TPC detector filled with 6 t liquid xenon
- By discarding the S1 signal, the S2-only analysis can lower its recoil-energy threshold by ~7 times
- This enables sensitivity to sub-GeV dark-matter candidates such as dark photons and axion-like particles
- Dominant instrumental backgrounds:
 - Delayed electrons
 - Cathode event
- Sub-dominant physics backgrounds:
 - CEνNS
 - o Electronic recoils
 - Neutrons

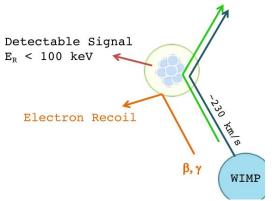


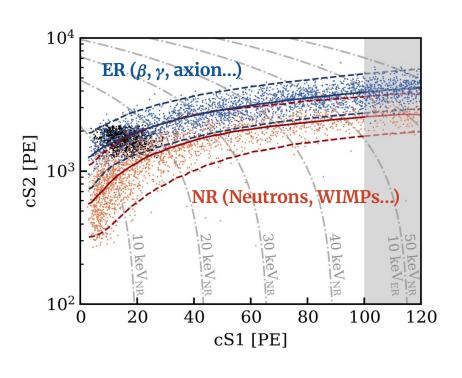
Back-up slides

How to identify NR (nuclear recoil)?

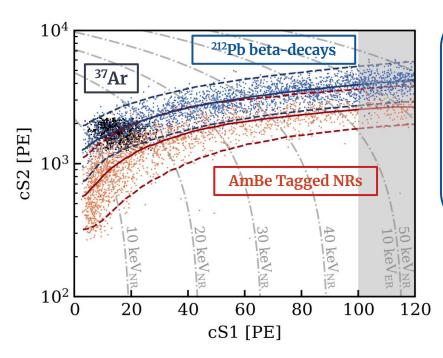
Using S2/S1 to discriminate electronic recoil (ER) and nuclear recoil (NR)







How to calibrate our detector?



Calibration for ER:

ERs from ²¹²Pb beta-decays from injected gaseous ²²⁰Rn:

- Continuous spectrum
- To define cS1 vs cS2 response for ER
- To validate cut acceptance

ERs from injected gaseous ³⁷Ar:

- mono-energetic at 2.8 keV
- To validate the low-energy ER response

Calibration for NR:

NRs from 241 AmBe neutron source:

- Tagged by a coincident gamma captured by neutron veto
- To define cS1 vs cS2 response for NR

How to identify the background?

ER background:

• Dominated by ²¹⁴Pb (a daughter of ²²²Rn) beta-decays

Surface background:

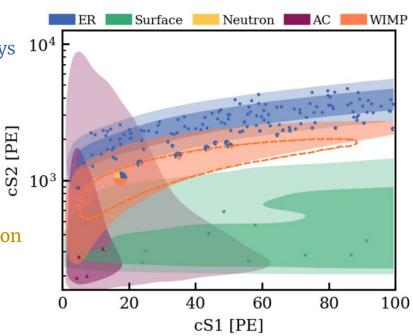
- beta decays of ²¹⁰Pb from TPC wall
- suppressed by fidicial volume cut

NR (neutron) background:

• Neutrons from spontaneous fission and (a,n) reaction

Accidental coincidence (AC) background:

• Random pairing of S1 and S2 lone signals



Upgrade from XENON1T to XENONnT

• Reduction of ER background:

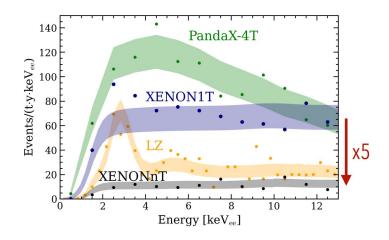
Major background: β emitter ²¹⁴Pb, a daughter of ²²²Rn

Rn distillation column ⇒Reduction of ER background by a factor of ~6

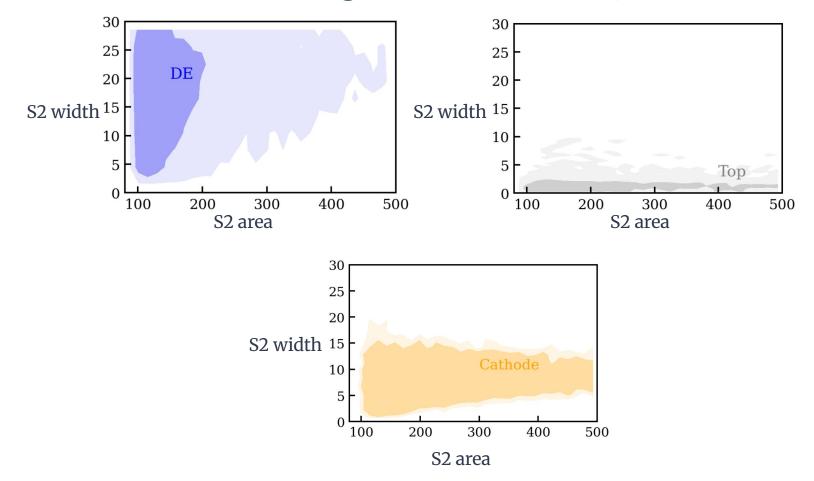
- New xenon purification system
- \Rightarrow Higher electron survival rate

| | Full drift time: | Electron lifetime: | Electron survival (@full drift length): |
|----|------------------|--------------------|--|
| 1T | 0.67 ms | 0.65 ms | 30 % |
| nT | 2.2 ms | ~15 ms | 86 % @ 15 ms |

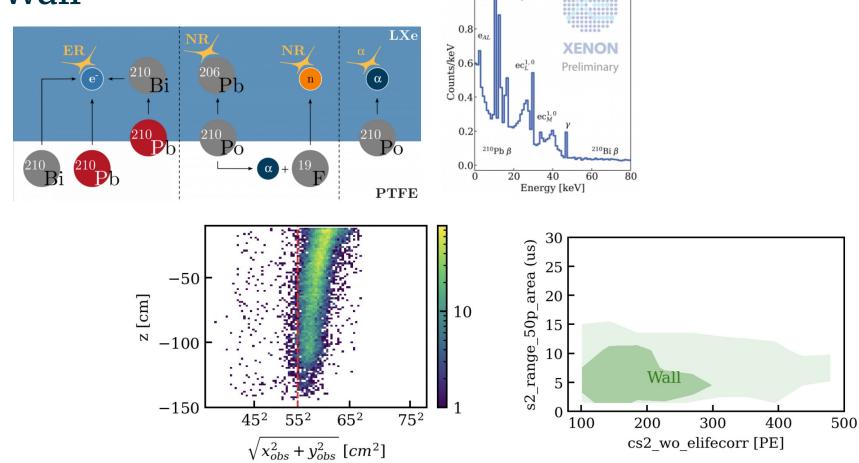




Major instrumental background in S2-only



Instrumental background: Wall



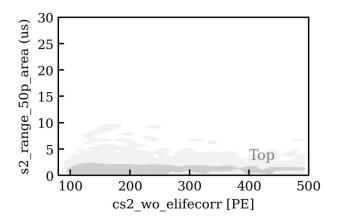
 $\times 10^4$

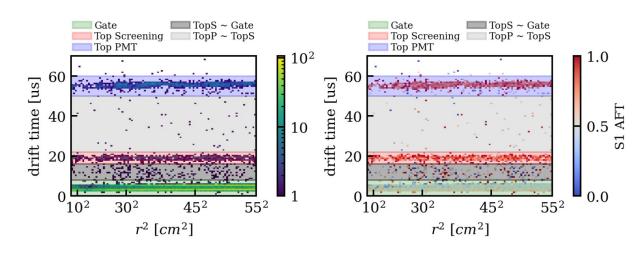
X-rays

1.0

Instrumental background:

Top electrodes





Physics background:

S2-only expected CEvNS events = 10 * 2 fold's

CEVNS

