

Leading-order term expansion for the Teukolsky equation on subextremal Kerr black holes

Pascal Millet

Ecole Polytechnique, CMLS

Conference of the GDR DynQua, CY Advanced Studies
Neuville-sur-Oise
01-02-2024

Classical field equations on a Kerr black hole

Equations of linearized gravity

$$D_{g_{M,a}} Ric(g) = 0$$

Equation of massless
neutrino

$$\nabla^{AA'} \phi_A = 0$$

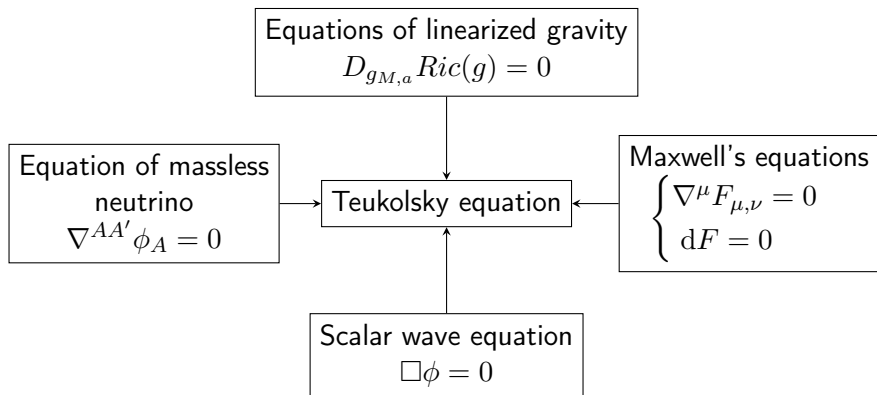
Maxwell's equations

$$\begin{cases} \nabla^\mu F_{\mu,\nu} = 0 \\ dF = 0 \end{cases}$$

Scalar wave equation

$$\square \phi = 0$$

Classical field equations on a Kerr black hole



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Here we focus on the case $\Lambda = 0$.

Kerr solutions

Kerr (1963): $\mathcal{M} = \mathbb{R}_t \times (r_+, +\infty) \times \mathbb{S}^2$, metric $g_{M,a}$

Model for a **rotating** black hole.

Subextremal: $|a| < M$.

$$g_{M,a} := \frac{\Delta_r - a^2 \sin^2 \theta}{\rho^2} dt^2 + \frac{4Mar \sin^2 \theta}{\rho^2} dt d\phi - \frac{\rho^2}{\Delta_r} dr^2 \\ - \rho^2 d\theta^2 - \frac{\sin^2 \theta}{\rho^2} ((a^2 + r^2)^2 - a^2 \Delta_r \sin^2 \theta) d\phi^2$$

$$\Delta_r := a^2 + r^2 - 2Mr$$

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- Contains **trapped** null geodesics.

Diagram of the Kerr spacetime

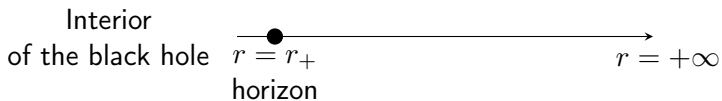


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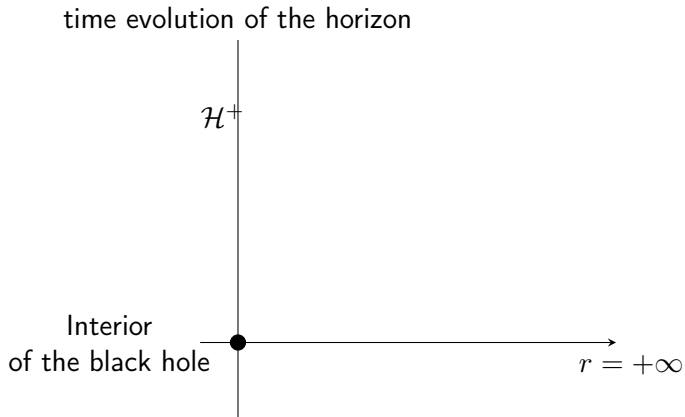


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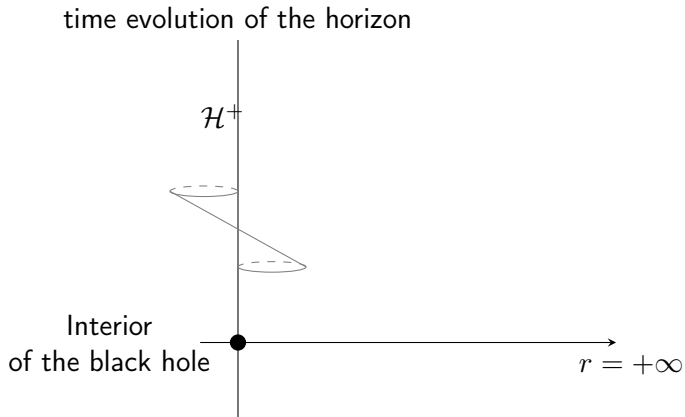


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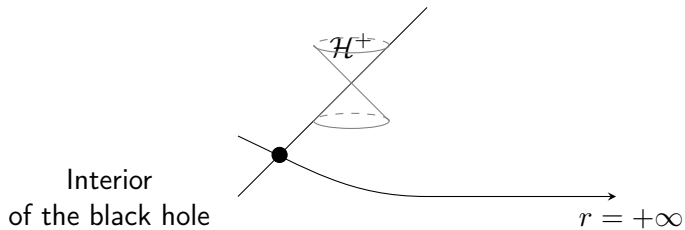


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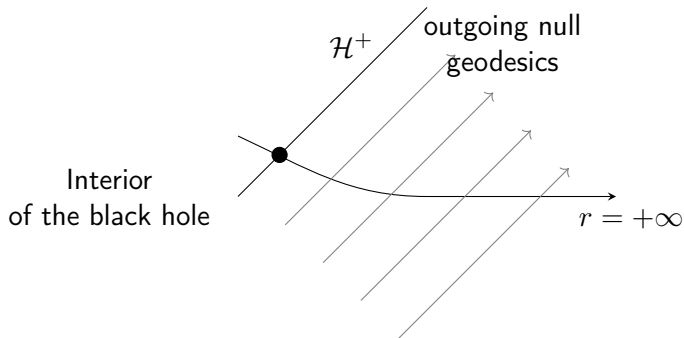


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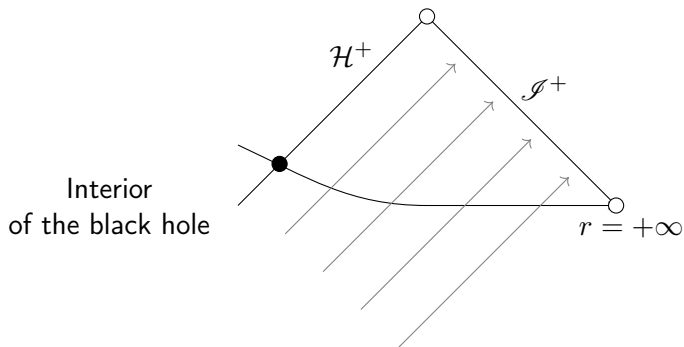
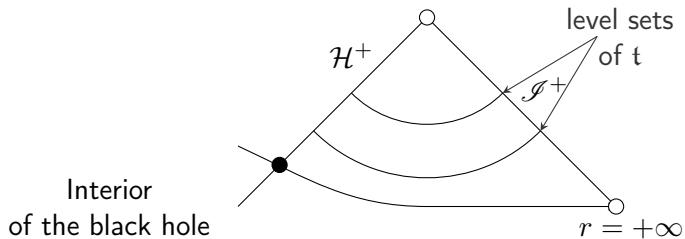


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$$\begin{aligned} \Psi_2 &= D_{g_{M,a}} W(\dot{g})(l, m, l, m) \\ \Psi_{-2} &= (r - ia \cos \theta)^4 D_{g_{M,a}} W(\dot{g})(n, \bar{m}, n, \bar{m}) \end{aligned}$$

Proposition (Teukolsky)

$D_{g_{M,a}} Ric(\dot{g}) = 0 \Rightarrow T_s \Psi_s = 0$ for $s = \pm 2$.

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In coordinates:

$$T_s = \rho^2 \square_{g_{M,a}} - \frac{2s(r-M)}{r^2 + a^2 \cos^2 \theta} \partial_r - 2s \left(\frac{a(r-M)}{\Delta_r} + \frac{i \cos \theta}{\sin^2 \theta} \right) \partial_\phi \\ - 2s \left(\frac{M(r^2 - a^2)}{\Delta_r} - r - ia \cos \theta \right) \partial_t + (s^2 \cot^2 \theta - s)$$

Theorem (M., 2023)

We consider a *subextremal* Kerr spacetime ($|a| < M$). We fix $s \in \frac{1}{2}\mathbb{Z}$. Let u_0, u_1 *smooth and compactly supported* on Σ_0 . The solution u of the Cauchy problem

$$\begin{cases} T_s u = 0 \\ u(t=0) = u_0 \\ \frac{\partial}{\partial t} u(t=0) = u_1 \end{cases}$$

satisfies:

$$|u(r, t, \theta, \phi) - \mathfrak{p}_{u_0, u_1}(r, t, \theta, \phi)| \leq Cr^{-1+t} t^{-2-|s|+s-\epsilon} \left(\frac{t}{r} + 1\right)^{-1-s-|s|}$$

where $\epsilon > 0$.

Related results on Teukolsky

Ma-Zhang ('21): case $|a| \ll M$, $s = \pm 1, \pm 2$.

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$$p = t^{-3-2|s|} \frac{(2|s| + 2) \left(\frac{t}{r}\right)^{2+|s|+s} + 2(|s| - s + 1) \left(\frac{t}{r}\right)^{1+|s|+s}}{\left(\frac{t}{r} + 2\right)^{2+|s|+s}} F_{u_0, u_1}$$

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where $F_{u_0, u_1}(r, \theta, \phi)$ can be expressed with hypergeometric functions and spin weighted spherical harmonics.

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- We can assume H^N regularity with N large instead of smooth.
- We get a precise **decay estimate** (without leading order term) when initial data only have inverse polynomial decay.

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- Inverse Fourier transform:

$$v(\mathbf{t}) = \int_{\mathbb{R}+i\mathbb{C}} e^{-i\sigma\mathbf{t}} R(\sigma)\hat{f}(\sigma) d\sigma$$

Formally

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- 2 Bound for $R(\sigma)$ when $|\sigma| \rightarrow +\infty$.
 - Semiclassical estimates near radial points Vasy ('11, '19)
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- 4 Regularity of $R(\sigma)$ near zero. Principal term in the development \Leftrightarrow highest order singularity (here at zero).

2: High frequency bound

$\|R(\sigma)\|_{\mathcal{L}(\mathcal{Y}, \mathcal{X})} \leq C$ when $|Re(\sigma)| \rightarrow +\infty$, $\Im(\sigma)$ bounded.
small parameter: $h = \frac{1}{|\sigma|}$, $\sigma = h^{-1}z$

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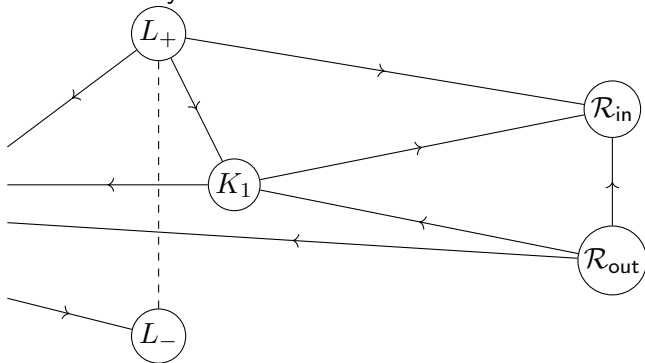
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$u = \sum_{i=1}^N \chi_i(x, hD)u$, χ_i microlocalizers.

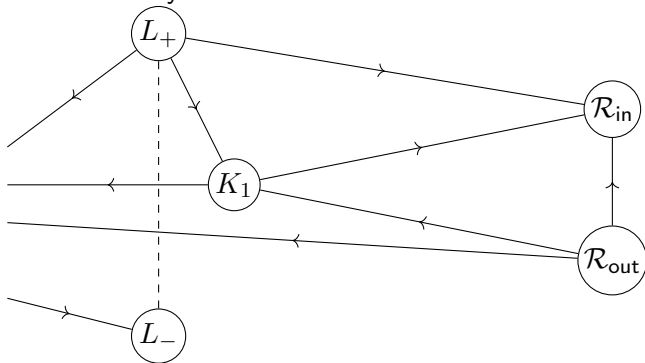
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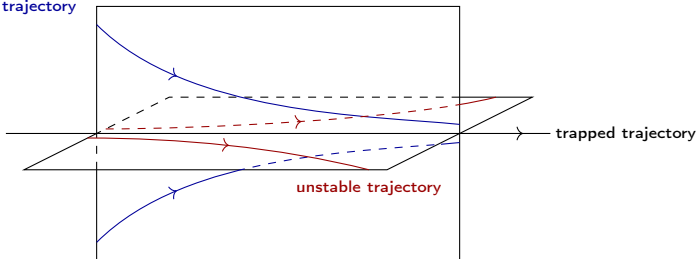
- Detailed study of the semiclassical Hamiltonian flow.



- Semiclassical version of the radial points estimates (Vasy '13, '19)

- Estimates at the normally hyperbolic trapped set (Wunsch-Zworski '14, Dyatlov '14)

stable trajectory



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- More precise result in the case of non compact support.
- Use the result to study Maxwell and linearized gravity (decay and scattering theory) in the case $|a| < M$.

Thank you for your attention !