



Finanțat de
Uniunea Europeană
NextGenerationEU



Planul Național
de Redresare și Reziliență

PNRR: Fonduri pentru România modernă și reformată!

Acceleration as a circular motion along an imaginary circle

Victor E. Ambruș¹, Maxim N. Chernodub^{2,1}

PLB 855 (2024) 138757

¹ West University of Timișoara, Romania

² Institut Denis Poisson, Tours, France

8th χ rality | 22nd July 2024



„PNRR. Finanțat de Uniunea Europeană – Următoarea Generație UE”

Motivation

Accelerating states:

- ▶ Unruh effect [W. G. Unruh, PRD **14** (1976) 870]
- ▶ Global thermodynamic equilibrium state [F. Becattini, PRD **97** (2018) 085013]
- ▶ Black hole evaporation
- ▶ Early universe

Why investigate?

- ▶ In rigid rotation, $\langle T^{\mu\nu} \rangle$ receives quantum corrections from ω and a .
- ▶ What is the effect of acceleration on strongly-interacting matter EoS?
- ▶ Lattice studies require imaginary time and suitable boundary conditions \Rightarrow what is the KMS relation giving these bcs?

Global equilibrium with constant acceleration a

- ▶ In global equilibrium, the temperature four-vector $\beta^\mu = u^\mu/T$ satisfies the Killing equation:

$$\partial_\mu \beta_\nu + \partial_\nu \beta_\mu = 0. \quad (1)$$

- ▶ For constant acceleration,

$$\beta^\mu \partial_\mu = \beta_T [(1 + az)\partial_t + at\partial_z], \quad (2)$$

where $\beta_T = 1/T$ represents the (constant) inverse temperature,

- ▶ Imposing $u^2 = 1$, we have

$$T(x) = \frac{1}{\beta_T} [(1 + az)^2 - (at)^2]^{-1/2}, \quad (3)$$

$$u^\mu \partial_\mu = \beta_T T(x) [(1 + az)\partial_t + at\partial_z], \quad (4)$$

$$a^\mu \partial_\mu = a\beta_T^2 T^2(x) [at\partial_t + (1 + az)\partial_z]. \quad (5)$$

- ▶ The state diverges at the Rindler horizon:

$$(1 + az)^2 - (at)^2 = 0, \quad z \geq -\frac{1}{a}. \quad (6)$$

- ▶ The proper thermal acceleration $\alpha^\mu = a^\mu/T(x)$ has constant magnitude:
 $\alpha = \sqrt{-\alpha^\mu \alpha_\mu} = a\beta_T.$

Density operator for global equilibrium

- ▶ The density operator reads

$$\hat{\rho} = \exp \left[- \int d\Sigma_\mu T^{\mu\nu} \beta_\nu \right] = e^{-b \cdot \hat{P} + \varpi : \hat{J} / 2}, \quad (7)$$

where $b^\mu = b^\mu = \beta_T \delta_0^\mu$ and $\varpi^\mu{}_\nu = \alpha(\delta_3^\mu g_{0\nu} - \delta_0^\mu g_{3\nu})$.

- ▶ Taking into account the property $\mathsf{T}(a)\Lambda\mathsf{T}(a)^{-1} = \mathsf{T}[(\mathbb{I} - \Lambda)(a)]\Lambda$, one can factorize $\hat{\rho}$ as

[Becattini et al, JHEP 02(2021)101]

$$\hat{\rho} = e^{-\tilde{b}(\varpi) \cdot \hat{P}} e^{\varpi : \hat{J} / 2}, \quad (8)$$

where

$$\tilde{b}(\varpi)^\mu = \sum_{k=0}^{\infty} \frac{i^k}{(k+1)!} (\varpi^\mu{}_{\nu_1} \varpi^{\nu_1}{}_{\nu_2} \cdots \varpi^{\nu_{k-1}}{}_{\nu_k}) b^{\nu_k} = B\delta_0^\mu + A\delta_3^\mu, \quad (9)$$

$$B = \frac{\sin \alpha}{a}, \quad A = \frac{i}{a}(1 - \cos \alpha),$$

such that $\hat{\rho} = e^{-B\hat{H} + A\hat{P}^z} e^{\alpha\hat{K}^z}$.

KMS relation for accelerating states

- ▶ Taking into account the properties

$$e^{i\tilde{b}\cdot\hat{P}}\hat{\Phi}(x)e^{-i\tilde{b}\cdot\hat{P}} = \hat{\Phi}(x + \tilde{b}), \quad \hat{\Lambda}\hat{\Phi}(x)\hat{\Lambda}^{-1} = D[\Lambda^{-1}]\hat{\Phi}(\Lambda x), \quad (10)$$

we arrive at $\hat{\rho}\hat{\Phi}(t, z)\hat{\rho}^{-1} = e^{-\alpha S^{0z}}\hat{\Phi}(\tilde{t}, \tilde{z})$, where

$$\begin{aligned} \tilde{t} &= \cos(\alpha)t + i\sin(\alpha)z + \frac{i}{a}\sin(\alpha), \\ \tilde{z} &= i\sin(\alpha)t + \cos(\alpha)z - \frac{1}{a}[1 - \cos(\alpha)]. \end{aligned} \quad (11)$$

- ▶ The KMS are formulated at the level of the Wightman functions:

$$G^+(x, x') = \langle \hat{\Phi}(x)\hat{\Phi}(x') \rangle, \quad G^-(x, x') = (-1)^{2s} \langle \hat{\Phi}(x')\hat{\Phi}(x) \rangle, \quad (12)$$

with $s = 0$ for scalars and $s = 1/2$ for Dirac fermions.

- ▶ At finite temperature and under acceleration, we derive the KMS relations:

$$G^+(x, x') = (-1)^{2s} e^{-\alpha S^{0z}} G^-(\tilde{t}, \tilde{z}; x'). \quad (13)$$

Euclidean two-point functions: rational acceleration

- ▶ The KMS relations transfer to the Euclidean two-point functions:

$$G_E(\tilde{\tau}, \tilde{z}; X') = (-1)^{2s} e^{-\alpha S^{0z}} G_E(\tau, z; X'), \quad (14)$$

which are solved formally by

$$G_E^{(\alpha)}(\tau, z; X') = \sum_{j=-\infty}^{\infty} (-1)^{2sj} e^{-j\alpha S^{0z}} G_E^{\text{vac}}(\tau_{(j)}, z_{(j)}; X'), \quad (15a)$$

where $G_E^{\text{vac}}(X, X')$ is the vacuum Euclidean propagator, while

$$\begin{aligned} \tau_{(j)} &= \tau \cos(j\alpha) - \frac{1}{a}(1 + az) \sin(j\alpha), \\ z_{(j)} &= \tau \sin(j\alpha) + \frac{1}{a}(1 + az) \cos(j\alpha) - \frac{1}{a}. \end{aligned} \quad (16)$$

- ▶ When $\alpha/2\pi = p/q \in \mathbb{Q}$, the sum over j in Eqs. (15) contains only q terms:

$$G_E^{(p,q)}(\tau, z; x') = \sum_{j=0}^{q-1} (-1)^{2sj} e^{-j\alpha S^{0z}} G_E^{\text{vac}}(\tau_{(j)}, z_{(j)}; x'). \quad (17)$$

Energy density

- ▶ Using

$$G_E^{\text{vac}}(X, X') = \frac{1}{4\pi^2 \Delta X^2}, \quad S_E^{\text{vac}}(X, X') = \gamma_\mu^E \partial_\mu G_E^{\text{vac}} \quad (18)$$

we find

(in agreement with [*Becattini*, JHEP, 202*])

$$\mathcal{E}_{\text{KG}}^\xi = \frac{3[\alpha T(x)]^4}{16\pi^2} [G_4(\alpha) + 4\xi G_2(\alpha)], \quad \mathcal{E}_D = \frac{3[\alpha T(x)]^4}{4\pi^2} S_4(\alpha),$$

$$G_n(\alpha) = \sum_{j=1}^{\infty} \frac{1}{[\sin(j\alpha/2)]^n}, \quad S_n(\alpha) = - \sum_{j=1}^{\infty} \frac{(-1)^j \cos(j\alpha/2)}{[\sin(j\alpha/2)]^n}.$$

- ▶ Since $e^{-q\alpha S^{0z}} = (-1)^p$, $S_E^{(p,q)}(\tau_{(q)}, z_{(q)}) = (-1)^{p+q} S_E^{p,q}(\tau, z)$, such that $S_E^{(p,q)} = 0$ when $p + q$ is odd.
- ▶ **For** $\alpha = p/q$: In this case, both G_n and S_n can be computed in closed form. We find:

$$\mathcal{E}_\xi^{(p,q)} = \frac{[\alpha T(x)]^4}{480\pi^2} (q^2 - 1)(q^2 + 11 + 60\xi),$$

$$\mathcal{E}_D^{(p,q)} = \frac{[\alpha T(x)]^4}{960\pi^2} (q^2 - 1)(7q^2 + 17) \frac{1 + (-1)^{p+q}}{2}. \quad (19)$$