

Acceleration as a circular motion along an imaginary circle

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Abstract

We describe a quantum fluid undergoing constant acceleration a in the grand canonical ensemble, in thermal equilibrium at finite inverse temperature β_T . Writing the action of the density operator $\hat{\rho}$ as a Poincaré transformation with imaginary parameters, we derive the Kubo-Martin-Schwinger (KMS) relation characterizing the two-point functions. The KMS relation sets boundary conditions for the Euclidean propagator, identifying points in the $\tau - z$ plane on a circle separated by an angle equal to the thermal acceleration $\alpha = a\beta_T$. When $\alpha/2\pi = p/q$ is a rational number, we find a fractalization of thermodynamics, similar to the case of states under imaginary rotation. **(Based on Ref. [1])**

1 Global thermodynamic equilibrium

- In kinetic theory, a fluid is in thermal equilibrium if $f_{\mathbf{k}} = [e^{\beta_{\mu}k^{\mu} - \alpha} + \varepsilon]^{-1}$, with $\varepsilon = \pm 1$ for Fermi-Dirac/Bose-Einstein statistics.
- Imposing the Boltzmann equation, $k^{\mu}\partial_{\mu}f_{\mathbf{k}} = C[f] = 0$, the temperature four-vector $\beta^{\mu} = u^{\mu}/T(x)$ satisfies the Killing equation, $\partial_{\mu}\beta_{\nu} + \partial_{\nu}\beta_{\mu} = 0$, while $\alpha = \mu(x)/T(x) = \text{const}$.
- β^{μ} can be written in terms of the 10 Killing vectors of Minkowski space [2]:

$$\beta^{\mu} = b^{\mu} + \varpi^{\mu}{}_{\nu}x^{\nu}, \quad (1)$$

with constant b^{μ} and $\varpi^{\mu\nu} = -\varpi^{\nu\mu}$, such that $\beta_{\mu}\beta^{\mu} = 1/T^2(x) > 0$.

- For **constant acceleration**: $\beta^{\mu}\partial_{\mu} = \beta_T[(1 + az)\partial_t + at\partial_z]$, $[\beta_T, a \equiv \text{const}]$
- The density operator corresponding to Eq. (1) reads [3, 4]

$$\hat{\rho} = \exp\left(-b_{\mu}\hat{P}^{\mu} + \frac{1}{2}\varpi_{\mu\nu}\hat{J}^{\mu\nu}\right) = e^{-\tilde{b}(\varpi)\cdot\hat{P}}e^{\varpi\cdot\hat{J}/2}, \quad (2)$$

where for simplicity, we set $\alpha = \mu = 0$ and $\hat{\rho}$ was factorized using

$$\tilde{b}^{\mu} = \sum_{k=0}^{\infty} \frac{i^k}{(k+1)!} (\varpi^{\mu}{}_{\nu_1}\varpi^{\nu_1}{}_{\nu_2}\dots\varpi^{\nu_{k-1}}{}_{\nu_k})b^{\nu_k}. \quad (3)$$

- For **constant acceleration**: $\tilde{b}^{\mu} = B\delta_0^{\mu} + A\delta_3^{\mu}$, $\varpi^{\mu\nu} = \alpha(-g^{\mu t}g^{\nu z} - g^{\mu z}g^{\nu t})$ and

$$\hat{\rho} = e^{-B\hat{H} + A\hat{P}^z}e^{\alpha\hat{K}^z}, \quad A = \frac{i}{a}(1 - \cos\alpha), \quad B = \frac{\sin\alpha}{a}, \quad (4)$$

with $\alpha = \sqrt{-\alpha^{\mu}\alpha_{\mu}} = a\beta_T = \text{const}$ and $\alpha^{\mu} = a^{\mu}/T(x)$ the thermal acceleration, while

$$a^{\mu}\partial_{\mu} = (u^{\nu}\partial_{\nu}u^{\mu})\partial_{\mu} = aT^2(x)\beta_T^2[at\partial_t + (1 + az)\partial_z]. \quad (5)$$

2 Quantum thermal KMS relation

- The quantum field operator $\hat{\Phi}(x)$ is covariant under Poincaré transformations:

$$e^{i\tilde{b}\cdot\hat{P}}\hat{\Phi}(x)e^{-i\tilde{b}\cdot\hat{P}} = \hat{\Phi}(x + \tilde{b}), \quad \hat{\Lambda}\hat{\Phi}(x)\hat{\Lambda}^{-1} = D[\Lambda^{-1}]\hat{\Phi}(\Lambda x), \quad (6)$$

where $\Lambda = e^{-\frac{i}{2}\varpi\cdot\mathcal{J}}$, $(\mathcal{J}^{\mu\nu})_{\alpha\beta} = i(\delta_{\alpha}^{\mu}\delta_{\beta}^{\nu} - \delta_{\beta}^{\mu}\delta_{\alpha}^{\nu})$ and $D[\Lambda]^{-1} = e^{\frac{i}{2}\varpi\cdot S}$ is the spin part of the inverse Lorentz transformation. $[S = 0 \text{ for the scalar field; and } S = \frac{i}{2}\gamma^{\beta}\gamma^{\alpha}\gamma^{\gamma} \text{ for the Dirac field}]$

- The density operator acts like a Poincaré transformation with imaginary parameters:
- For **constant acceleration**: $\hat{\rho}\hat{\Phi}(t, \varphi)\hat{\rho}^{-1} = e^{-\alpha S^{0z}}\hat{\Phi}(\tilde{t}, \tilde{z})$, where $[S^{0z} = \frac{i}{2}\gamma^0\gamma^z]$

$$\tilde{t} = \cos(\alpha)t + i\sin(\alpha)z + \frac{i}{a}\sin(\alpha), \quad \tilde{z} = i\sin(\alpha)t + \cos(\alpha)z - \frac{1}{a}[1 - \cos(\alpha)]. \quad (7)$$

- We consider now the scalar/Dirac Wightman functions:

$$\begin{aligned} G^+(x, x') &= \langle \hat{\Phi}(x)\hat{\Phi}(x') \rangle, & S^+(x, x') &= \langle \hat{\Psi}(x)\hat{\Psi}(x') \rangle, \\ G^-(x, x') &= \langle \hat{\Phi}(x')\hat{\Phi}(x) \rangle, & S^-(x, x') &= -\langle \hat{\Psi}(x')\hat{\Psi}(x) \rangle, \end{aligned} \quad (8)$$

with $\langle \hat{A} \rangle = Z^{-1}\text{Tr}(\hat{\rho}\hat{A})$ and $Z = \text{Tr}(\hat{\rho})$. Using $\langle \hat{A} \rangle = \text{Tr}(\hat{\rho}\hat{B}\hat{\rho}\hat{A}\hat{\rho}^{-1})$ leads to the KMS relations:

- For **constant acceleration** [1]:

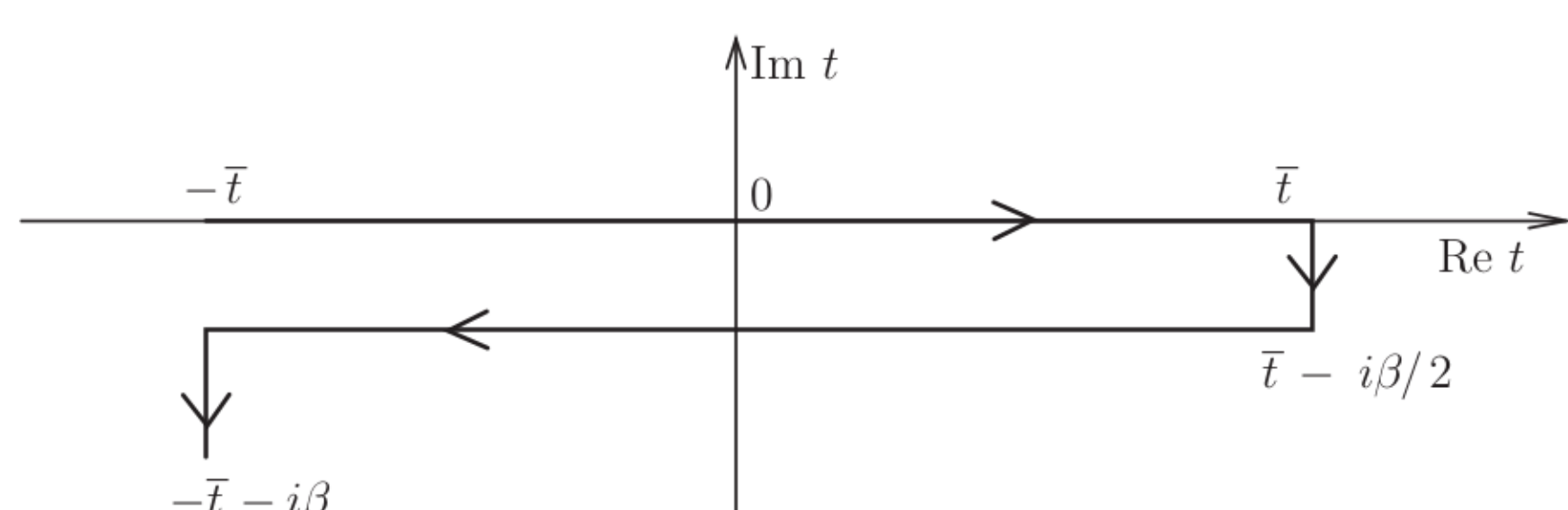
$$G^+(t, z; x') = G^-(\tilde{t}, \tilde{z}; x'), \quad S^+(t, \varphi; x') = -e^{-\alpha S^{0z}}S^-(t + i\beta_T, \varphi + i\beta_T\Omega; x'). \quad (9)$$

3 Feynman propagator under constant acceleration

- We now consider the Feynman two-point function,

$$G_F(x, x') = \Theta_C(x, x')G^+(x, x') + \Theta_C(x', x)G^-(x, x'),$$

where $\Theta_C(x, x')$ is the causal step function along the thermal contour.



- The KMS relations imply:

$$G_F(\tilde{t}, \tilde{z}; x') = G_F(t, z; x'), \quad S_F(\tilde{t}, \tilde{z}; x') = -e^{-\alpha S^{0z}}S_F(t, z; x'). \quad (10)$$

- Keeping these relations with respect to the imaginary time $\tau = it$, a formal solution can be written in terms of the Euclidean vacuum propagators,

$$\begin{aligned} G_E^{(\alpha)}(\tau, z; X') &= \sum_{j=-\infty}^{\infty} G_E^{\text{vac}}(\tau_j, z_j; X'), \\ S_E^{(\alpha)}(\tau, z; X') &= \sum_{j=-\infty}^{\infty} (-1)^j e^{-j\alpha S^{0z}} S_E^{\text{vac}}(\tau_j, z_j; X'). \end{aligned}$$

with $j \in \mathbb{Z}$ and

$$\begin{aligned} \tau_j &= \tau \cos(j\alpha) - \frac{1}{a}(1 + az) \sin(j\alpha), \\ z_j &= \tau \sin(j\alpha) + \frac{1}{a}(1 + az) \cos(j\alpha) - \frac{1}{a}. \end{aligned}$$

- Using $G_E^{\text{vac}}(X, X') = 1/(4\pi^2\Delta X^2)$ and $S_E^{\text{vac}}(X, X') = \gamma_{\mu}^E\partial_{\mu}G_E^{\text{vac}}$, we find the energy density $\mathcal{E} = u_{\mu}T^{\mu\nu}u_{\nu}$ to be (in agreement with Refs. [3, 4])

$$\mathcal{E}_{\text{KG}}^{\xi} = \frac{3[\alpha T(x)]^4}{16\pi^2} [G_4(\alpha) + 4\xi G_2(\alpha)], \quad \mathcal{E}_D = \frac{3[\alpha T(x)]^4}{4\pi^2} S_4(\alpha), \quad (11)$$

where $G_n(\alpha) = \sum_{j=1}^{\infty} [\sin(j\alpha/2)]^{-n}$ and $S_n(\alpha) = -\sum_{j=1}^{\infty} (-1)^j \cos(j\alpha/2) / [\sin(j\alpha/2)]^n$.

4 Rational acceleration

- In the case when $\alpha/2\pi = p/q$ is a rational number, the contour closes on itself after q iterations:

$$\tau_{(q)} = \tau, \quad z_{(q)} = z,$$

such that

$$\begin{aligned} G_E^{(p,q)}(\tau, z) &= \sum_{j=0}^{q-1} G_E^{\text{vac}}(\tau_j, z_j), \\ S_E^{(p,q)}(\tau, z) &= \sum_{j=0}^{q-1} (-1)^j e^{-j\alpha S^{0z}} \\ &\quad \times S_E^{\text{vac}}(\tau_j, z_j). \end{aligned}$$

- Since $e^{-q\alpha S^{0z}} = (-1)^p$, $S_E^{(p,q)}(\tau_{(q)}, z_{(q)}) = (-1)^{p+q} S_E^{(p,q)}(\tau, z)$, such that $S_E^{(p,q)} = 0$ when $p + q$ is odd.
- In this case, both G_n and S_n can be computed in closed form. We find:

$$\begin{aligned} \mathcal{E}_{\xi}^{(p,q)} &= \frac{[\alpha T(x)]^4}{480\pi^2} (q^2 - 1)(q^2 + 11 + 60\xi), \\ \mathcal{E}_D^{(p,q)} &= \frac{[\alpha T(x)]^4}{960\pi^2} (q^2 - 1)(7q^2 + 17) \frac{1 + (-1)^{p+q}}{2}. \end{aligned} \quad (12)$$

References

- [1] V. E. Ambruș, M. N. Chernodub, Phys. Lett. B **855** (2024) 138757.
- [2] C. Cercignani, G. M. Kremer, *The Relativistic Boltzmann Equation: Theory and Applications* (Springer, 2002).
- [3] F. Becattini, M. Buzzegoli, A. Palermo, JHEP **02** (2021) 101.
- [4] A. Palermo, M. Buzzegoli, F. Becattini, JHEP **10** (2021) 077.

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