# The magnetic dissipative effect on quark-gluon plasma

# direct photon production and hyperon local spin polarization

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with Li Yan









# **O** The magnetic field in HIC and weak magnetic effect in QGP **O** The direct photon $v_2$ and the weak magnetic emission **O** The spin polarization and the weak magnetic polarization O Summary and outlook

# The magnetic field in the heavy-ion collisions



T. Bowman and J. Abramowitz/Brookhaven National Laboratory

- o There must be a B field generated.
- o orientated out of plane
- o Extremely strong initially.

$$eB/m_{\pi}^2 \sim \begin{cases} O(1) & \text{RHIC} \\ O(10) & \text{LHC} \end{cases}$$

 $m_{\pi}^2 \approx 10^{17}$  Gauss



L. Yan and X.-G. Huang (2021), 2104.00831 vacuum Deng W T, Huang X G. Phys. Rev. C, 2012, 85: 044907.

- o The B field decays dramatically
- o B field during the QGP expansion is weak.

$$|B(\tau_0)| \sim 10^{-3} |B(0)| \ll m_\pi^2$$

A. Huang et.al (2022),2212.08579.

J.-J. Zhang, et.al, Phys. Rev. Res. 4, 033138 (2022)



# The magnetic effect in QGP: weak vs strong

A slightly redistribution of the



A weak B field:

o The QCD matter dynamics is merely affected: scattering process, transport coefficients... The magnetic field can be safely viewed as the perturbations of hydro background

Huang, Zhao and Zhuang, 2208.01407

## The dissipation due to EM field: w/o spin

• The dissipative correction from the Chapman-Enskog expansion:  $p^{\mu}\partial_{\mu}f + qF^{\mu\nu}p_{\mu}\frac{1}{\dot{\epsilon}}$ at the leading order of  $\frac{|eB|}{T^2}$  $\delta f_{\rm EM} \sim \tau_R$ 

o Landau matching: dissipative correction in conserved current

$$J_{a,\mu} = eQ_a n_a u^{\mu} + eQ_a N_{a,\mu} \qquad N_{a,\mu} = \int \frac{d^3 p}{(2\pi)^3 E_p} p^{\mu} \delta f_{a,EM}$$

$$J_{\mu} = \sum_{a} J_{a,\mu}$$

 $|eB| \ll T^2$ 

$$\frac{\partial}{\partial p^{\nu}} f = C[f] \sim \frac{f - n_{\rm eq}}{\tau_R}$$

Vlasov term

$$_{R}qF^{\mu\nu}p_{\mu}rac{\partial}{\partial p^{
u}}n_{\mathrm{eq}}$$

$$= \sigma_{el} F_{\mu\nu} u^{\nu} \equiv \sigma_{el} E_{\mu\nu}$$



## The dissipation due to EM field: with spin

### • The Chapman-Enskog expansion with spin d.o.f:

$$p^{\mu}\partial_{\mu}\mathscr{F} + QF^{\mu\nu}p_{\mu}\frac{\partial\mathscr{F}}{\partial p^{\nu}} = -\mathscr{C}[\mathscr{F}] = -(p \cdot u)\frac{\mathscr{F} - \mathscr{F}_{eq}}{\tau_{R}}$$

$$\mathcal{F}_{eq} = \frac{1}{2m}$$

$$\delta \mathscr{F}_{\rm EM} = -\frac{\bar{\tau}}{T} Q F^{\mu\nu} p_{\mu} \frac{\partial}{\partial p^{\nu}} \mathscr{F}_{\rm eq} = \frac{1}{2m} \overline{U}(p) Y(x,p) U(x,p)$$

Y(x,

### • The spin averaged quark distribution function reduces to

$$\frac{1}{2} \operatorname{tr}_2 \mathscr{F}_{eq} = n_{eq}$$

F. Becattini et al., Annals of Physics 338 (2013) 32–49

Effects of dissipation S. Bhadury's talk on Friday

 $\overline{U}(p)X(x,p)U(p)$ 

$$p) \equiv -\frac{\bar{\tau}}{T}QF^{\mu\nu}p_{\mu}\frac{\partial X}{\partial p^{\nu}} = \frac{\bar{\tau}}{T}QF^{\mu\nu}p_{\mu}\beta_{\nu}e^{\beta\cdot p}X^{2}\exp\left(-\frac{1}{2}\omega_{\alpha\beta}\Sigma^{\alpha\beta}\right)$$

$$\frac{1}{2} \operatorname{tr}_2 \delta \mathscr{F}_{\rm EM} = \delta f_{\rm EM}$$





# **Direct** photon $v_2$ : theoretical expectation

### **o** Direct photons: all sources except hadron decay.

• Theory expects smaller  $v_2$  of direct photons than hadrons





# **Direct photon** v<sub>2</sub>: **Experimental results**



PHENIX, PRL (2012)





# **Experiments** vs theory

## Hydrodynamical models



• "Not too much of a puzzle left for yields." [K. Reygers, Quark Matter 2022 plenary talk] • The present models are being challenged. A. Adare et al. (PHENIX), Phys. Rev. C94, 064901 (2016) 9

## Transport calculations

### Fireball model



# **Direct photon** $v_2$ : the most updated calculations

o Pre-equilibrium dynamics (KoMPost)
o Chemical equilibration in QGP
o NNLO pQCD for prompt photons
o Dissipation corrections from shear and bulk

$$f_q = n_q + \delta f$$

The  $v_2$  of direct photon is still **under-predicted**.

*The gap*  $\sim 0.05 - 0.10^{-6}$ 

J-F Paquet, et,al Phys. Rev. C93, 044906 (2016) C. Gale, J.-F. Paquet, B. Schenke, and C. Shen, Phys.Rev. C 105, 014909(2022)





# The weak magnetic photon emission

• Small angle approximation





• No significant difference beyond the small angle approximation

J.Sun et.al, in progress 11





# The source of momentum anisotropy

of the fireball



[STAR collaboration, PRL 101, 252301 (2008)] Gursoy, Dima Kharzeev and Rajagopal, PRC (2014)

## • The coupling effect between the weak magnetic field and the longitudinal dynamics







# EBE hydro v<sub>2</sub>: RHIC

### AuAu@200GeV



- A realistic simulation: Trento3D + MUSIC (1,000 events each centrality)
- o "All" source of direct  $\gamma$  + weak magnetic emissions
- A dimensionless parameter  $\rho$  tuned to cover the data
- The photon elliptic flow can be enhanced significantly and confront the experiment data with the weak magnetic emission.



# **EBE hydro yield: RHIC**



o The tuned  $\rho$  for  $v_2^{\gamma}$  is used to calculate the photon yields. • The increased yields  $\sim 10\%$ -20%, small and acceptable.

# EBE hydro v<sub>3</sub>: RHIC



• The significant triangle flow increment demonstrates the non-trivial weak magnetic effect furthermore.



# The Global spin polarization

## o the most vortical system $\omega = (9 \pm 1) \times 10^{21} s^{-1}$



• Consistent with the hydrodynamic prediction

• How the orbit angler momentum is converted to the spin of particles





### Along the beam direction



STAR 2023 PhysRevLett.131.202301

Alzhrani S, Ryu S, Shen C. Physical Review C, 2022, 106(1): 014905.

- 0
- o Sign observed against naive hydro expectation using thermal vorticity.

In thermal equilibrium: thermal vorticity driven by initial geometry, leading to spin polarization.





## **Beyond global equilibrium: thermal shear coupling**



- The sign can be flipped by the SIP(BBP).
- With SIP, the third-order modulation is greater than the second-order one.
- The centrality dependence can not be reproduced with SIP.

S. Y. F. Liu and Y. Yin, JHEP 07, 188 (2021) F. Becattini, M. Buzzegoli, and A. Palermo, 18 Phys. Lett. B 820, 136519 (2021), 18 arXiv:2103.10917



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A. Bzdak, V. Skokov, 1111.1949



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## The weak magnetic polarization

### **Dissipative terms introduced by the weak B**

$$\mathcal{F} = \mathcal{F}_{eq} + \delta \mathcal{F}_{EM} = \frac{1}{2m} \bar{U}(p)(X+Y)U(p)$$

• Spin tensor:

$$s^{\lambda,\rho\sigma}(x) = \frac{1}{2} \int \frac{d^3 \boldsymbol{p}}{2p^0} \operatorname{tr}_2\left(\mathscr{F}\overline{U}(p)\{\gamma^\lambda, \Sigma^{\rho\sigma}\}U(p)\right) = \int \frac{d^3 \boldsymbol{p}}{2p^0} \left[p^\lambda \Theta^{\rho\sigma} + p^\rho \Theta^{\sigma\lambda} + p^\sigma \Theta^{\lambda\rho}\right]$$

$$\Theta^{\mu\nu}(x) \equiv \operatorname{tr}_2\left[(X+Y)\Sigma^{\mu\nu}\right] = \left[n_F(1-n_F) + (1-2n_F)\delta f_{\mathrm{EM}}\right]\omega^{\mu\nu}$$

 $P_{\mu}(x,p) = -\frac{1}{2\mathrm{tr}_2 f} \epsilon_{\mu\rho\sigma\tau} \frac{ds^{0}}{d^3\mu}$ o Polarization vector

$$\begin{split} \langle P_{\mu} \rangle &= -\frac{1}{4} \frac{p^{\tau}}{m} \epsilon_{\mu\rho\sigma\tau} \frac{\int d\Sigma \cdot p \; n_F (1 - n_F) \bar{\omega}^{\rho\sigma} + (1 - 2n_F) \delta f_{\rm EM} \bar{\omega}^{\rho\sigma}}{\int d\Sigma \cdot p \; {\rm tr} f} \\ &\approx -\frac{1}{8} \frac{p^{\tau}}{m} \epsilon_{\mu\rho\sigma\tau} \frac{\int d\Sigma \cdot p \; (n_F + \delta f_{\rm EM}) \bar{\omega}^{\rho\sigma}}{\int d\Sigma \cdot p \; (n_F + \delta f_{\rm EM})} \end{split}$$

F. Becattini et al., Annals of Physics 338 (2013) 32–49

$$\frac{0,\rho\tau}{^{3}p}\frac{p^{\tau}}{m} = -\frac{1}{2\mathrm{tr}X}\epsilon_{\mu\rho\sigma\tau}\Theta^{\rho\sigma}\frac{p^{\tau}}{m}$$



# The extracted sin modulation: eB dependence



- o Both modulations get increased monotonically with the B field introduced
- The ordering that the 2nd modulation > 3rd one, is found with the magnetic field.

**ally** with the B field introduced one, is found with the magnetic field.



### $\omega^{th} + \omega^{SIP(BBP)}$



- o The 2nd harmonics experimental data is used to extract the B field.
- o The extracted B field is applied to calculate the 3rd harmonics.
- o The 2nd and 3rd harmonics centrality trend can be well reproduced.







# The momentum dependence



- Our theoretical calculation gives consistent description of the  $p_T$  dependence.
- The deviations at large  $p_T$ , where hydrodynamics becomes invalid, shouldn't be a surprise.



• At low  $p_T$  region, 2nd>3rd, this ordering indicates the presence of a weak magnetic field.



# The time averaged B strength



 $\sim 0.1 m_{\pi}^2$ 

- o Weak B field in QGP
- o The extracted field strength grows as centrality increases.
- 0 It is too weak to induce the splitting between  $\Lambda$  and  $\bar{\Lambda}$  global polarization, as well, the photon



OThe elliptic and triangle flow of direct photon both get significant increments, which confronts the experimental data. The sign of  $P_{\Lambda}^{z}(\phi)$  is flipped and the centrality dependence are reproduced.

unchanged whatever the B field strength is.

Possible observables witnessing the novel effect:

Arxiv:2406.10041

Non-trivial coupling effect between the weak magnetic field and the longitudinal dynamics of the fireball!

- As a benchmark, if there is no rapidity-odd  $v_1$ , the  $v_2^{\gamma}$  and  $P_z(\phi)$  remains
- The polarization of di-leptons? The  $v_1$  splitting of mesons and baryons? The spin polarization in pA system?



# Thank you for your attention!

$$\begin{split} E_p \frac{d^3 \bar{N}}{d^3 \mathbf{p}} &= \int_V \bar{\mathcal{R}}^{\gamma}(P, X) = \bar{v}_0 (1 + 2\bar{v}_2 \cos 2\phi_p) \qquad E_p \frac{d^3 N_{\rm EM}}{d^3 \mathbf{p}} = \int_V \mathcal{R}_{\rm EM}^{\gamma}(P, X) = v_0^{\rm EM} (1 + 2v_2^{\rm EM}) \\ v_0^{\gamma} &= \bar{v}_0 + v_0^{\rm EM} , \quad v_2^{\gamma} = \frac{\bar{v}_2 \bar{v}_0 + v_2^{\rm EM} v_0^{\rm EM}}{\bar{v}_0 + v_0^{\rm EM}} \end{split}$$

## • Bjorken analysis for illustration

For background medium: 
$$n_{eq} = A_0(\tau, \eta_s, p_T, Y) + A_1(\tau, \eta_s, p_T, Y) \cos \phi_p$$
  
 $f_{EM} \propto QB_y \frac{\tau_R}{T} \frac{\sinh \eta_s}{\cosh(y - \eta_s)} (A_0 + A_1 \cos \phi_p) \cos \phi_p$  This  $\cos \phi$  is from weak magnetic field.  
 $= QB_y \frac{\tau_R}{T} \frac{\sinh \eta_s}{\cosh(y - \eta_s)} \left[ \frac{A_1}{2} + A_0 \cos \phi + \frac{A_1}{2} \cos 2\phi \right]$   
Rapidity-odd! Must be Rapidity-odd

Rapidity-oud: Must be Kapidity-oud 26

# Back up





# Back up



## The sum of quark and anti-quark contribution

 $[...\cos\phi] + ...\cos 2\phi + ...] \cos \phi$ 



**A Rapidity-odd** v<sub>1</sub> splitting has been experimentally measured!

 $v_2^{EM} \sim 0.5$ 



# The magnetic field profile

$$f_{\rm EM} = \frac{c}{8\alpha_{\rm EM}} \frac{\sigma_{\rm el} n_{\rm eq} (1 - n_{\rm eq})}{T^3 p \cdot u} e Q_f F^{\mu\nu} p_\mu u_\nu$$

o Electrical conductivity: LO pQCD evaluation (AMY). o  $\eta_s$  dependence is retained as in vacuum and the time averaged B field  $e\bar{B}$  is extracted.

$$\Gamma(\eta) = \frac{1}{(b^2/4 + \gamma^2 \tau_0^2 (\sinh \eta_s + v \cosh \eta_s)^2)^{3/2}} + \frac{1}{(b^2/4 + \gamma^2 \tau_0^2 (\sinh \eta_s - v \cosh \eta_s)^2)^{3/2}}$$

 $eB_v$ 

PRC 92 011901, PRC 96 044912. K. Hattori and X. Huang, 1609.00747 JETSCAPE framework, arxiv 1903.07706

$$= \overline{eB_y} \frac{\Gamma(\eta)}{\Gamma(0)}$$





 $\varepsilon^{\mu\rho\sigma\tau} \frac{1}{E} \hat{t}_{\rho} \xi_{\sigma\lambda} p^{\lambda} p_{\tau}$ 

SIP(BBP):

SIP(LY): 
$$\varepsilon^{\mu\rho\sigma\tau} \frac{1}{E} u_{\rho} \xi_{\sigma\lambda} p_{\perp}^{\lambda} p_{\tau} \qquad p_{\perp}^{\lambda} = p^{\lambda} - (u \cdot p) u^{\lambda}$$

$$\overline{eB_y} = 0.1m_\pi^2$$

 $\tau_0 = 0.4 \text{ fm}$ 

$$\xi^{\mu\nu} \equiv \frac{1}{2} \left[ \partial^{\mu} \left( \frac{u^{\nu}}{T} \right) + \partial^{\nu} \left( \frac{u^{\mu}}{T} \right) \right]$$

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